

Euclid Synergy with Rubin Observatory

Isobel Hook

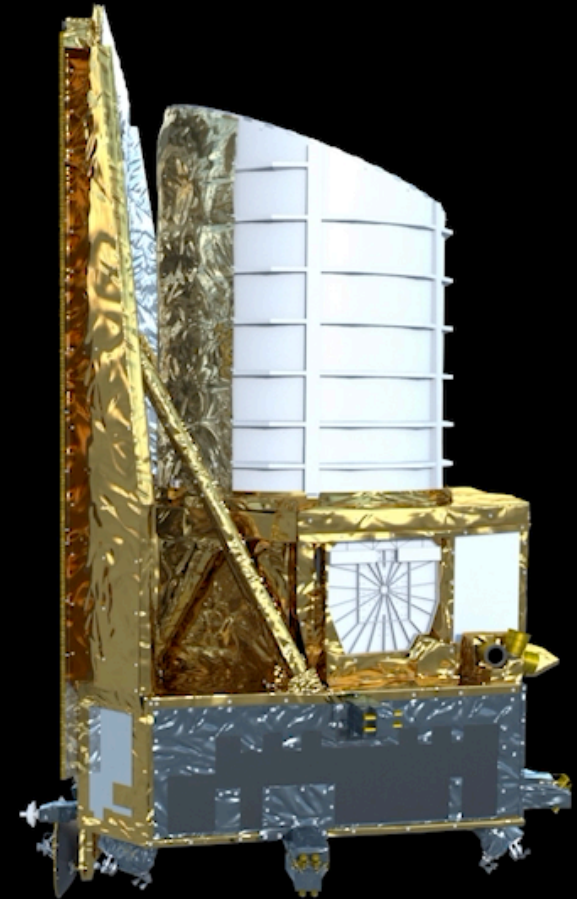
Lancaster University



Euclid

- ESA mission with focus on cosmology
 - Primary science: WL and BAO surveys
 - Huge legacy science potential
- 1.2 m telescope
- Launch planned for July – December 2022
- Nominal mission 6 years with possible 5-year extension
- Two wide-field instruments: VIS and NISP

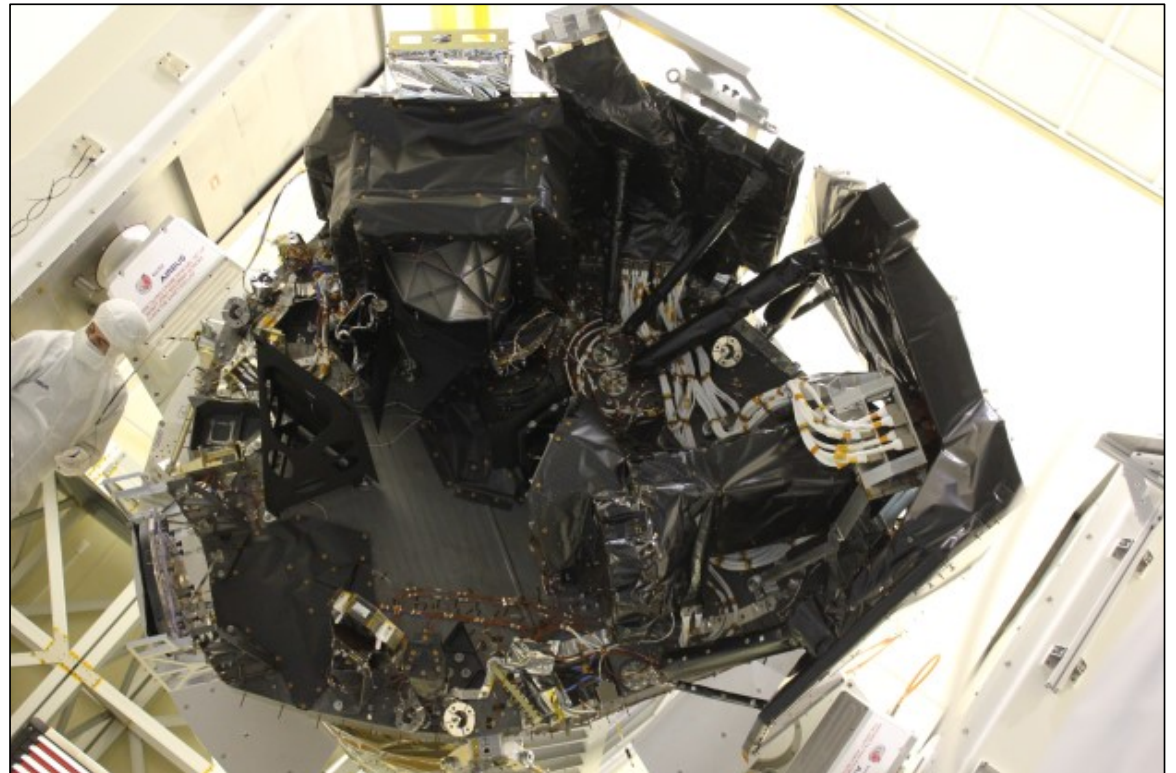
Euclid operating mode – credit: ESA/ATG medialab



Euclid Specifications

Euclid definition study report (“Red Book”), Laureijs et al 2011

- Wide survey 15,000 sq deg
- Wide survey depths
 - VIS imaging 10σ depth: 24.5
 - Y,J,H imaging 5σ depth: 24 mag (AB)
 - NIR spectroscopy at $R\sim 250$: unresolved line flux 3.5σ depth: 3×10^{-16} erg $\text{cm}^{-2}\text{s}^{-1}$
- Deep survey 40 sq deg
 - 2 magnitudes deeper
- PSF sampled with 0.1” pixels (VIS), 0.3” pixels (NISP)



Both instruments integrated onto the payload module of the spacecraft.
Image Credit: Airbus Defence and Space – Toulouse, Dec 2020

Science synergies with Rubin/LSST

(what does Euclid bring?)

Cosmology and Dark Energy

- Weak lensing
 - superior shape measurement and deblending
 - Note that Euclid *requires* optical photometry for phot-z and PSF colour
- BAO spectroscopic redshifts
 - Photo-z calibration
- Other cosmological probes
 - Clusters masses (evolution with z)
 - SNe Ia (host galaxies)

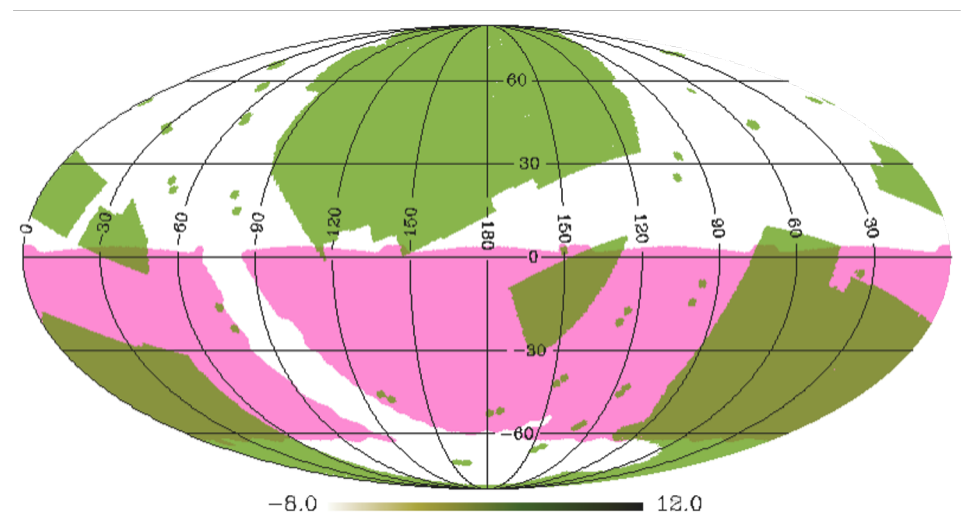
Legacy Science

- Solar System:
 - NIR colours , (sub-visit) astrometric shifts, spectra
- Milky Way and Local Group
 - star/galaxy separation, combined sensitivity to LSB features, e.g. stellar streams
- Galaxy evolution
 - galaxy stellar masses to $z \sim 3.5$, very high-z galaxies, morphology, spectra
- Transients:
 - IR detections, host galaxy properties

See e.g. Jain et al 2015, Rhodes et al 2017

Survey overlap is key

- Strong case for maximising survey overlap
- Euclid responses to LSST survey optimisation process
 - e.g. [Tri-Agency WG, Capak et al 2019](#)
- But note important publication policies: need Derived Data Products (DDP) process!

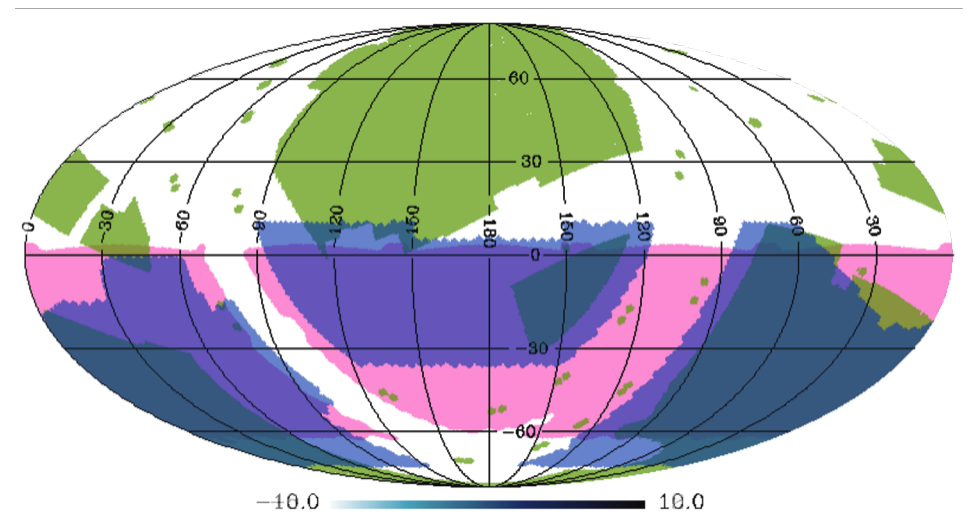


LSST baseline Nov 2018

Euclid ~early 2020

Survey overlap is key

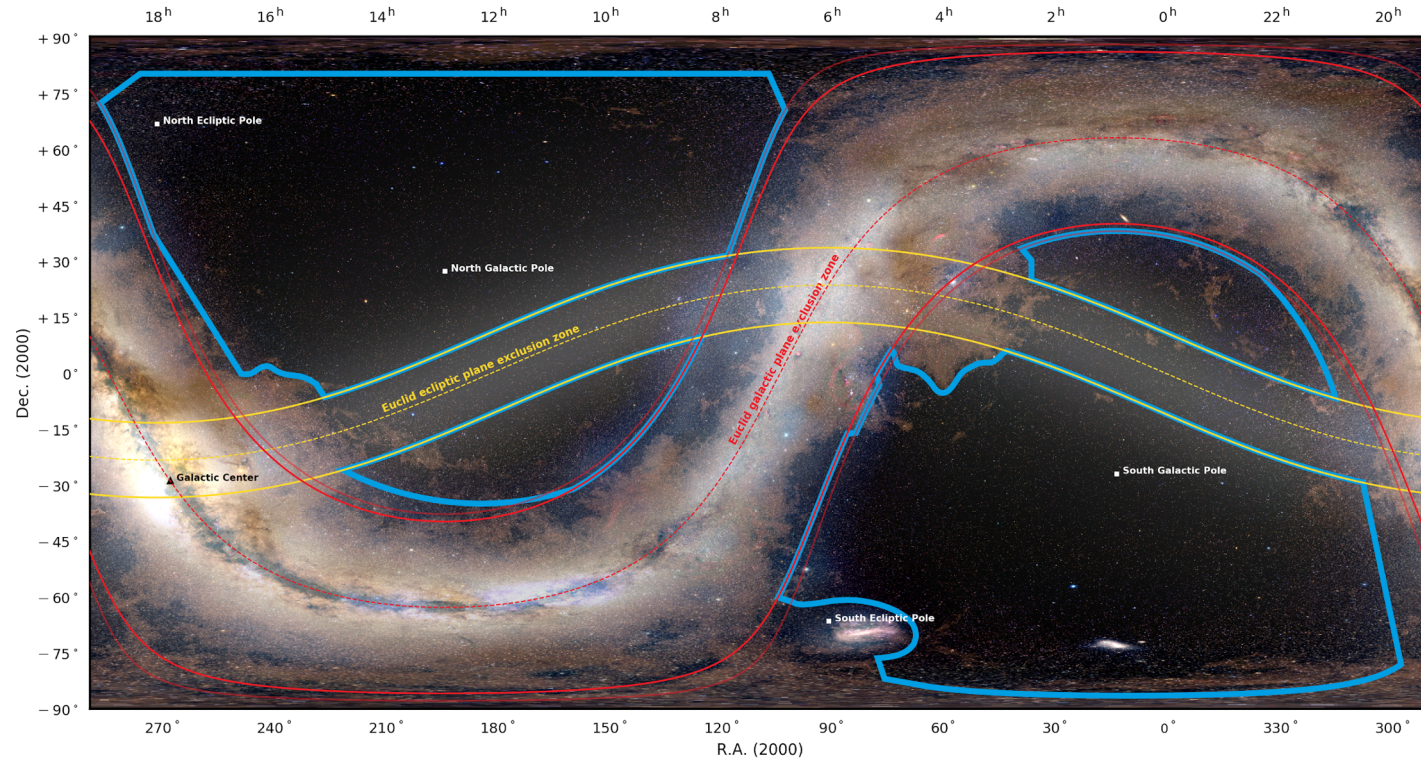
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LSST + 4MOST extragalactic overlap area $\sim 13\,000$ sq deg
LSST + 4MOST + Euclid overlap area ~ 6100 sq deg

Euclid Wide Survey : the region of interest

Slides from Jean-Charles Cuillandre



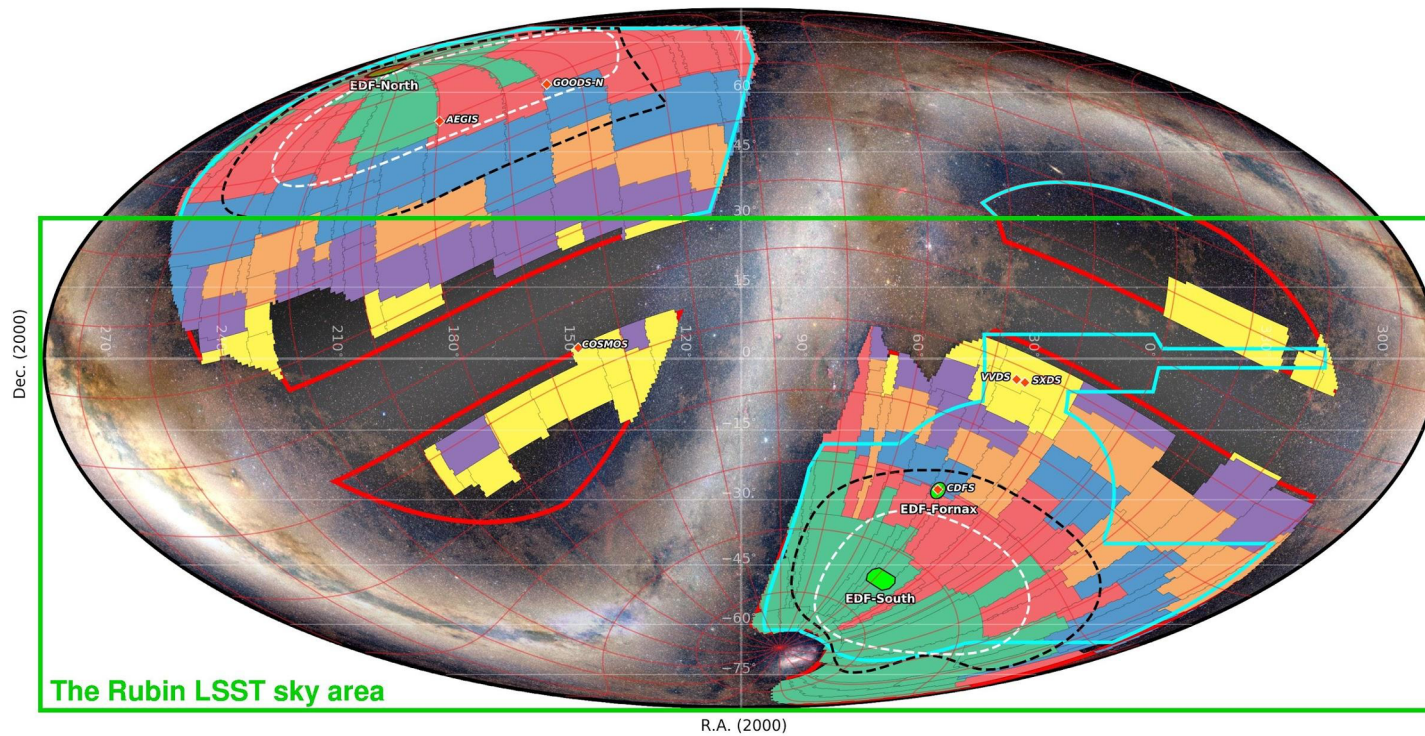
The Euclid Wide Survey based on ecliptic & galactic latitude thresholds + upper limits on stellar density & extinction (Gaia/Planck)

- ▭ Euclid Wide Survey region of interest : 17 Kdeg.² compliant with a 15 Kdeg.² survey
- ▭ Ecliptic plane [zodiacal light background] : +/- 10 deg. ecliptic latitude exclusion zone
- ▭ Galactic plane [stellar contamination] : +/- 23&25 deg. galactic latitude exclusion zone



Background image: Euclid Consortium / Planck Collaboration / A. Mellinger

The Euclid Wide Survey and the Rubin LSST sky area



The Rubin LSST sky area

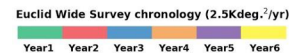
65% of the Euclid Region of Interest (17,354 square degrees)

RSD 2020c ECTile realization of a Euclid Wide Survey within the 17 Kdeg.² RoI : 14,668 deg.² over 6 years in 216 patches

Red outline: Euclid Wide Survey Region of Interest (RoI) : 17 Kdeg.² compliant with a 15 Kdeg.² survey

Black/White dashed outline: Best 2600 deg.² (black) and 1300 deg.² (white) SNR areas per galactic cap

Green outline: Euclid Deep Fields (EDF, from north to south): 10+10+23 deg.²

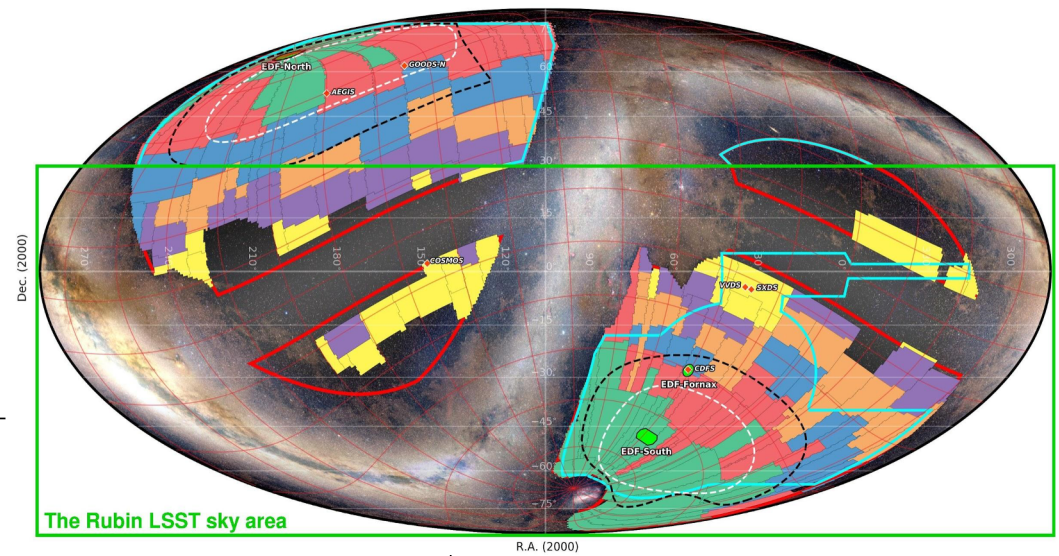


Background image: Euclid Consortium / Planck Collaboration / A. Mellinger

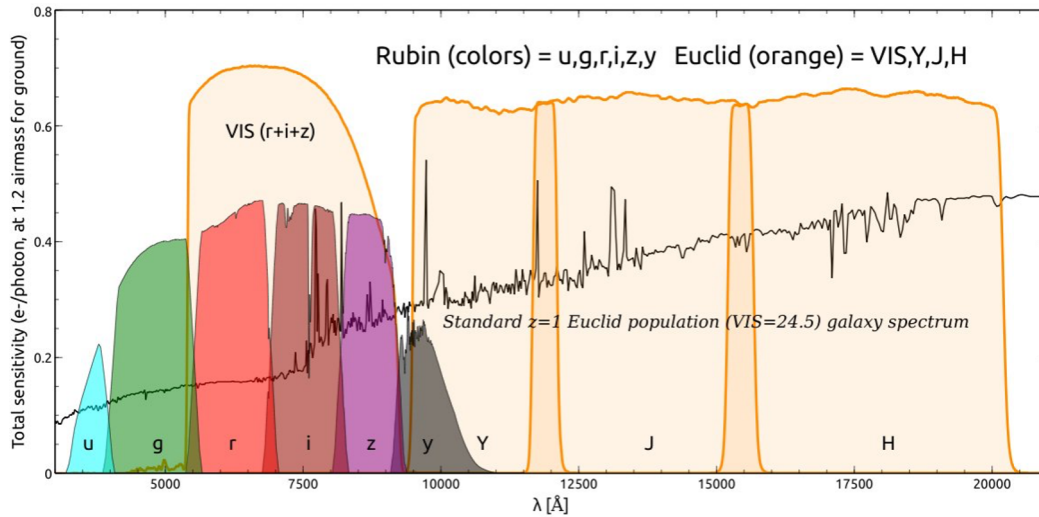
Euclid data releases

Assuming September 2022 Launch

- Jan 2023: End of commissioning
- Nov 2023: Q1
- Nov 2024: DR1 (2500 deg²) = year1 (green)
- Sep 2025: Q2
- Sep 2026: DR2 (7500 deg²) = year 1 to 3 (green+red+blue)
- Sep 2027: Q3
- Sep 2028: Q4 (TBC)
- Sep 2029: DR3 (15000 deg²) = year 1 to 6 (all colored areas)



Euclid performance as-built : deeper by 0.5 magnitude



Euclid (EOL) and Rubin (2018) Filter response curves

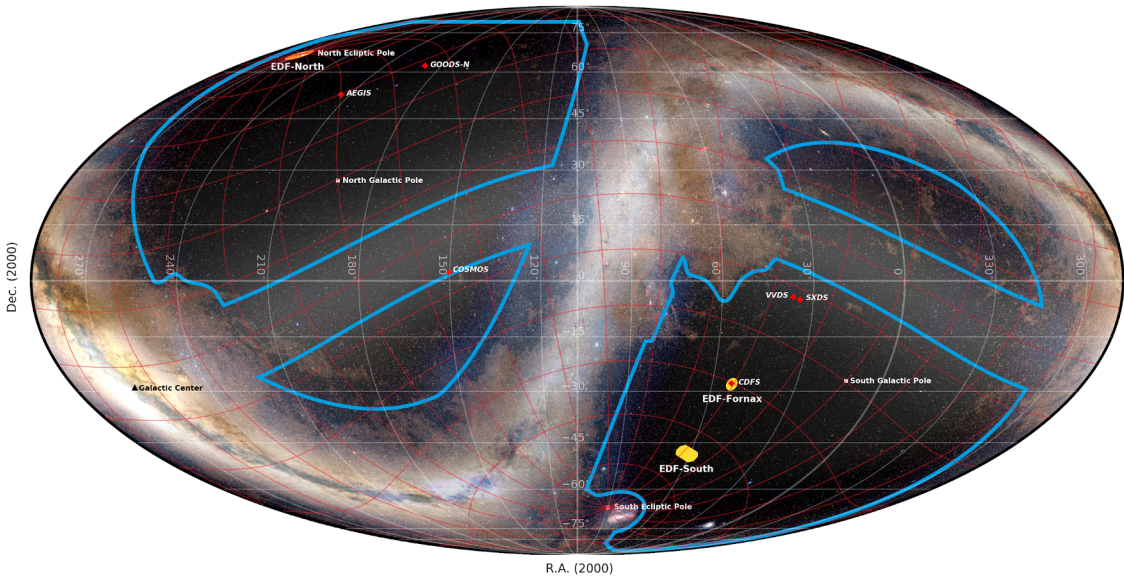
Euclid median SNR performance over the RoI for a VIS=24.5 extended source and a NISP Y, J, H 24th mag. star.
(Euclid Deep Survey: 2 mag fainter)

	<i>VIS</i>	<i>Y</i>	<i>J</i>	<i>H</i>	<i>S</i>
Minimum SNR	10.0	5.0	5.7	5.7	3.2
Median SNR	15.9	6.5	7.8	7.2	4.5
Maximum SNR	19.8	7.8	9.0	8.5	6.6
Median depth [AB mag]	26.2	24.3	24.5	24.4	—

Point source 5σ depth

Euclid Deep Survey : Rubin overlap on 2 fields (33 sq. deg.)

Location of the 3 Euclid Deep Fields

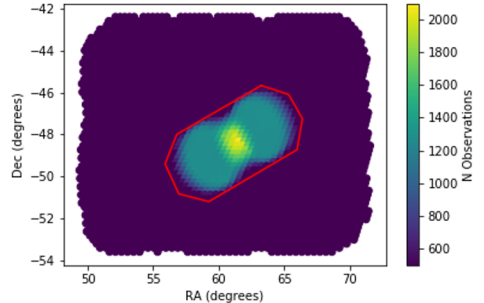


The Euclid Wide Survey (EWS) with the Euclid Deep Survey (EDF) and the deep Euclid Calibration Fields [Mollweide Celestial]

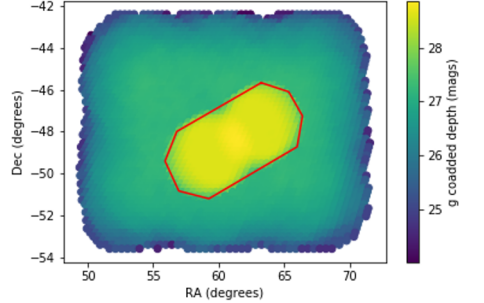
- ▭ Euclid Wide Survey region of interest : 17 Kdeg.² compliant with a 15 Kdeg.² survey
- ▭ Euclid Deep Fields : North=10 deg.², Fornax=10 deg.², South=23 deg.²
- ◆ Euclid deep calibration fields marker (diamond not to scale)



Background image: Euclid Consortium / Planck Collaboration / A. Mellinger



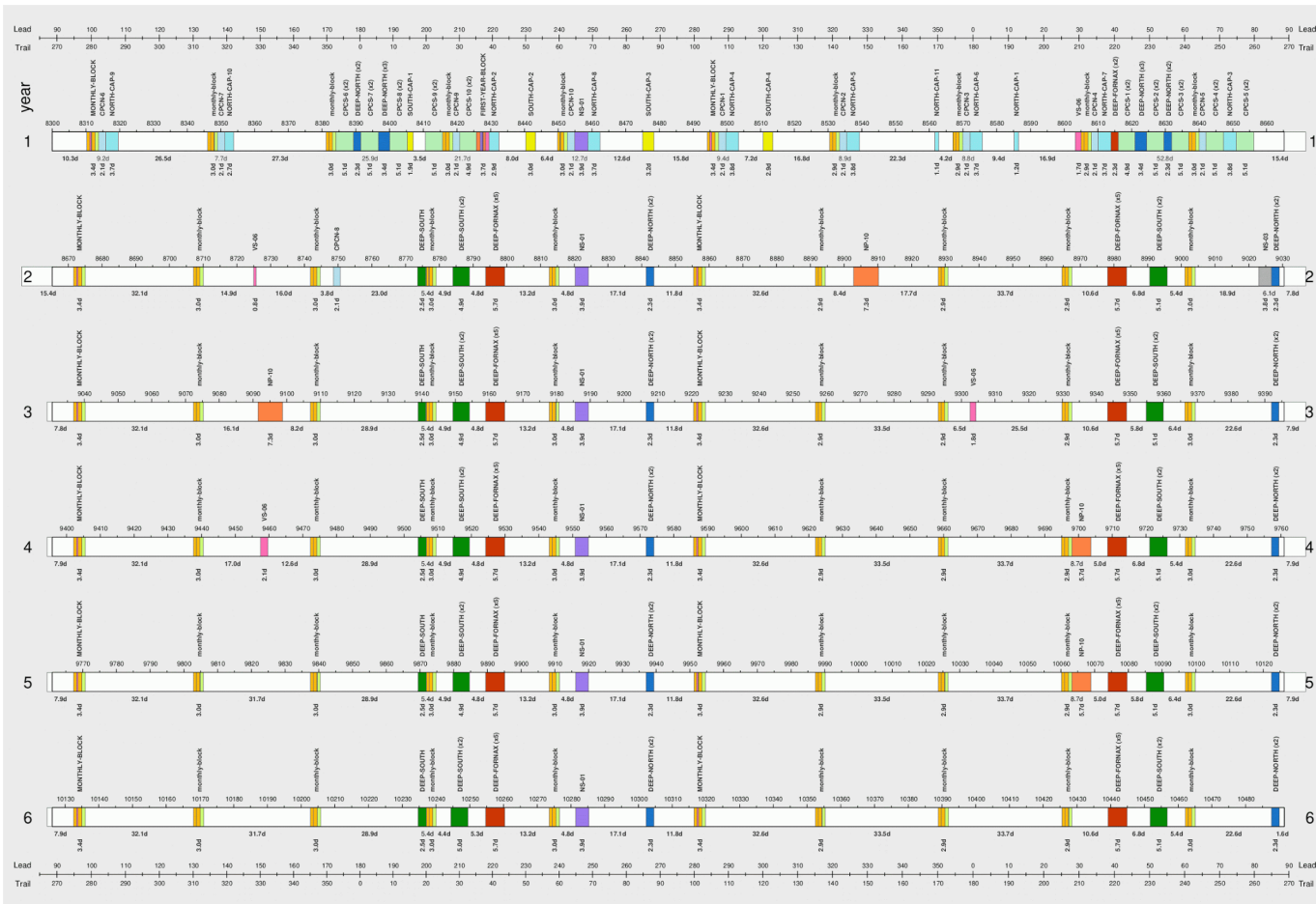
Two Rubin Deep Drilling Fields at half-depth on EDF-South



... or a uniform depth on EDF-South with an optimal Rubin dithering

EDF-South Rubin dithering (P. Yoachim)

Euclid Deep Survey : cadencing through the 6-year mission



Priority in Year 1 (2023) goes to EDF-North while EDF-Fornax (CDFs) and EDF-South reach full cadence in Year 2 (2024).

Depending on the launch date (currently September 2022), we will have a solid idea when EDF-Fornax and EDF-South ought to be observed.

The observing windows on the Euclid Deep Fields are 4 to 6-day long.

Plot from a 2020 ECSURV simulation

Conclusion

- Work ongoing to optimise survey overlap with LSST in space and time
- Science cases drive this
- Derived Data Products process is crucial!

