



# 4MOST – 4m Multi-Object Spectroscopic Telescope

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[www.4MOST.eu](http://www.4MOST.eu)



# 4MOST: Wide-field, high-multiplex optical spectroscopic survey facility for ESO

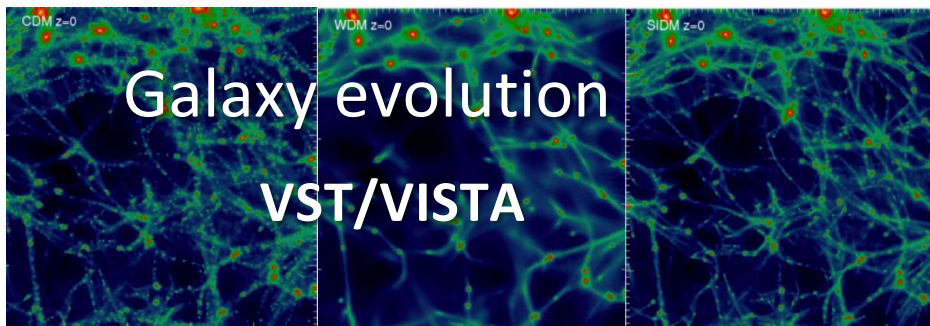


- **Status:**
  - ESO Council approved 4MOST in June 2015, ESO contract signed
  - Final Design Review in 2019, 4MOST is in full construction phase
  - Operations start on 4m **VISTA** telescope Q3-2023 (at least 2 x 5-year Surveys)
- **Survey facility:**
  - 4MOST Consortium delivers instrument, science operations, data products, science (70% of observing time)
  - All-sky 5 x 2 year public surveys in parallel, with yearly PUBLIC data releases
  - Key surveys organized by consortium in coordination with community
  - *Additional surveys from ESO community and Chile through ESO peer-review applications*



# Science Themes

4  
MOST



# Ten Consortium Surveys



No	Survey Name	Survey (Co-)PI
S1	Milky Way Halo LR Survey	Irwin (IoA), Starkenburg (RuG)
S2	Milky Way Halo HR Survey	Christlieb (ZAH)
S3	Milky Way Disk and Bulge LR Survey	Chiappini (AIP), Minchev (AIP)
S4	Milky Way Disk and Bulge HR Survey	Bensby (LU), Bergemann (MPIA)
S5	Galaxy Clusters Survey	Comparat (MPE)
S6	AGN Survey	Merloni (MPE)
S7	Galaxy Evolution Survey (WAVES)	Driver (USW), Liske (HHU)
S8	Cosmology Redshift Survey	Richard (CRAL), Kneib (EPFL)
S9	Magellanic Clouds Survey	Cioni (AIP)
S10	Time-Domain Extragalactic Survey (TiDES)	Sullivan (Southampton)



# Science Themes

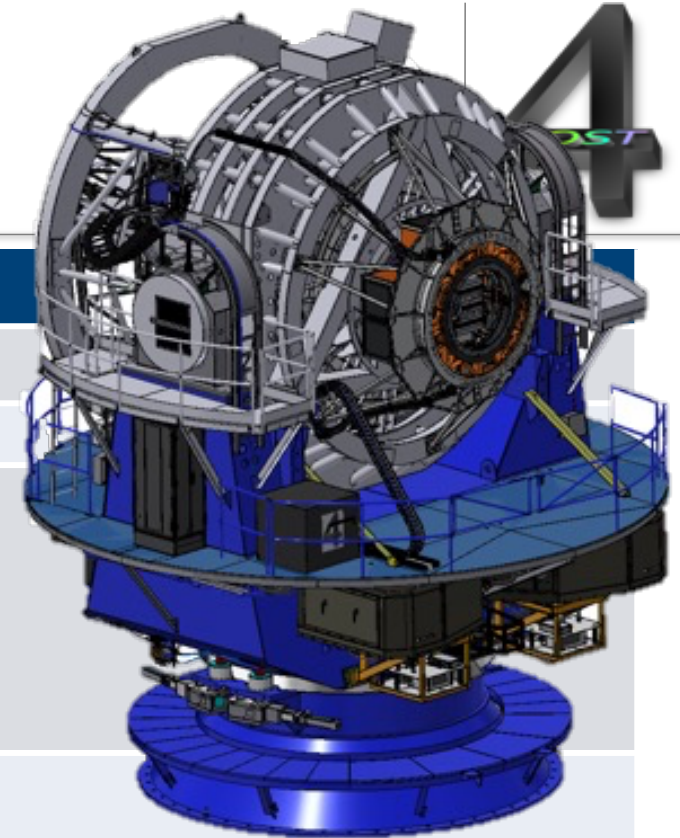


See also ESO Messenger ([2019Msng.175](#)) articles for more details

<https://www.eso.org/sci/publications/messenger/toc.html?v=175&m=Mar&y=19>

- Cosmology Survey: 10 million galaxy redshifts
- WAVES survey: 1.5 million galaxy redshifts
- eRosita X-ray Cluster survey: 50,000 cluster redshifts and velocity dispersions
- Supermassive Black Hole formation and evolution: 1.0 million AGN redshifts
- Distance Scale and Transient Universe: 200k SN (incl host galaxy) redshifts
- First Stars, Chemical Evolution of Milky Way: 2million high resolution spectra
- Merger, Accretion and Dynamical evolution of the Milky Way: 1million low(medium) resolution spectra

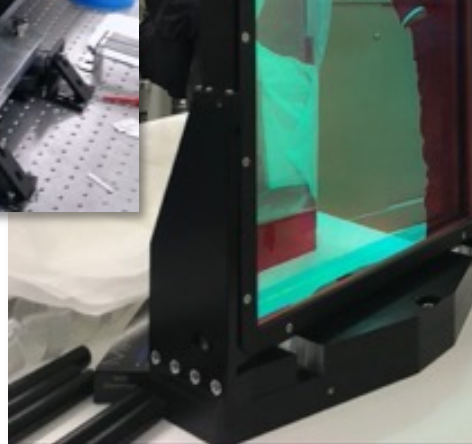
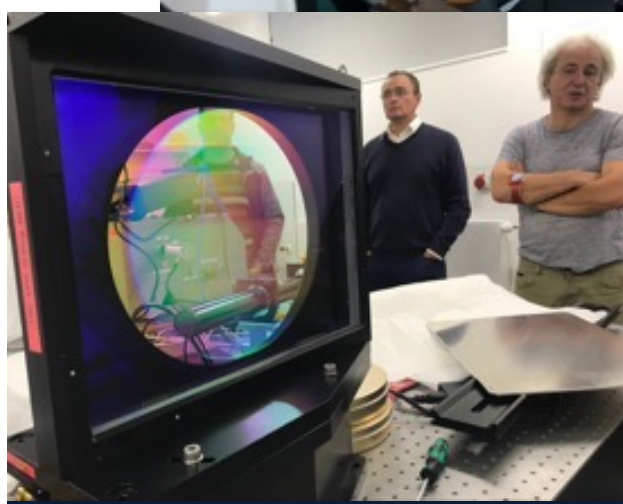
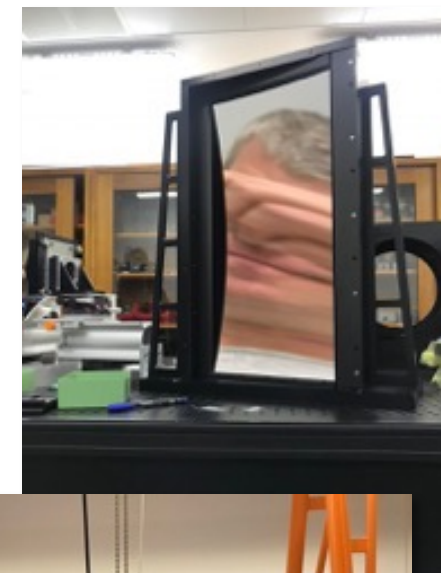
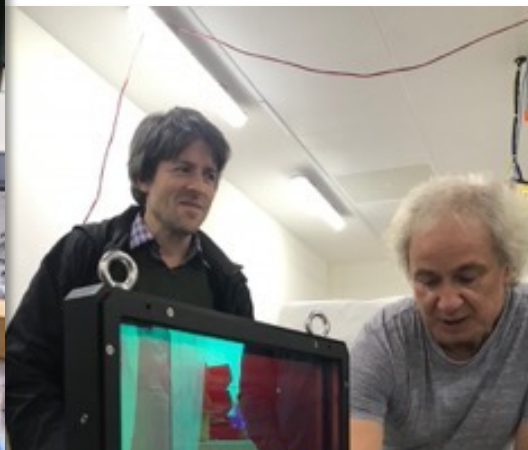
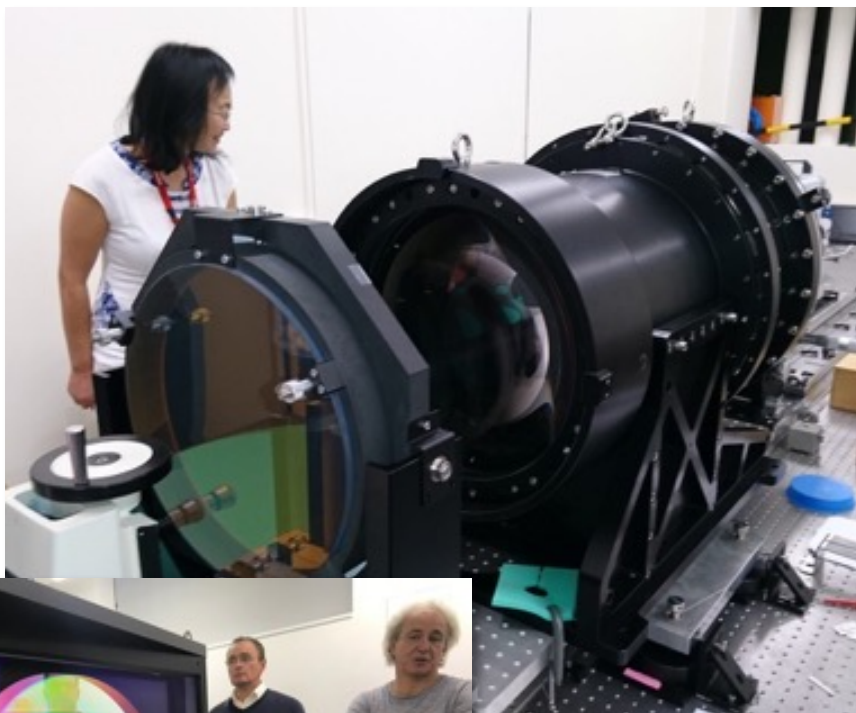
# Instrument Specification



Specification	Design value
Field-of-View (hexagon)	$\sim 4.2 \text{ degree}^2 (\varnothing > 2.6^\circ)$
Multiplex fiber positioner	2436
2 Low/Medium Resolution (LR) Spectrographs	R $\sim$ 4000–7500
Fibres	812 fibres (2x)
Passband	370–950 nm
Velocity accuracy	< 1 km/s
Spectral sampling (pixels/FWHM)	> 2.8 pixels
High Resolution (HR) Spectrograph	R $\sim$ 20,000
Fibres	812 fibres
Passband	392.6–435.5 nm, 516–573 nm, 610–679 nm
Velocity accuracy	< 1 km/s
Spectral sampling (pixels/FWHM)	> 2.56 pixels
# of fibers in $\varnothing=2'$ circle	>3
Fibre diameter	$\varnothing=1.42 \text{ arcsec}$

# High Resolution Spectrograph (HRS)

4  
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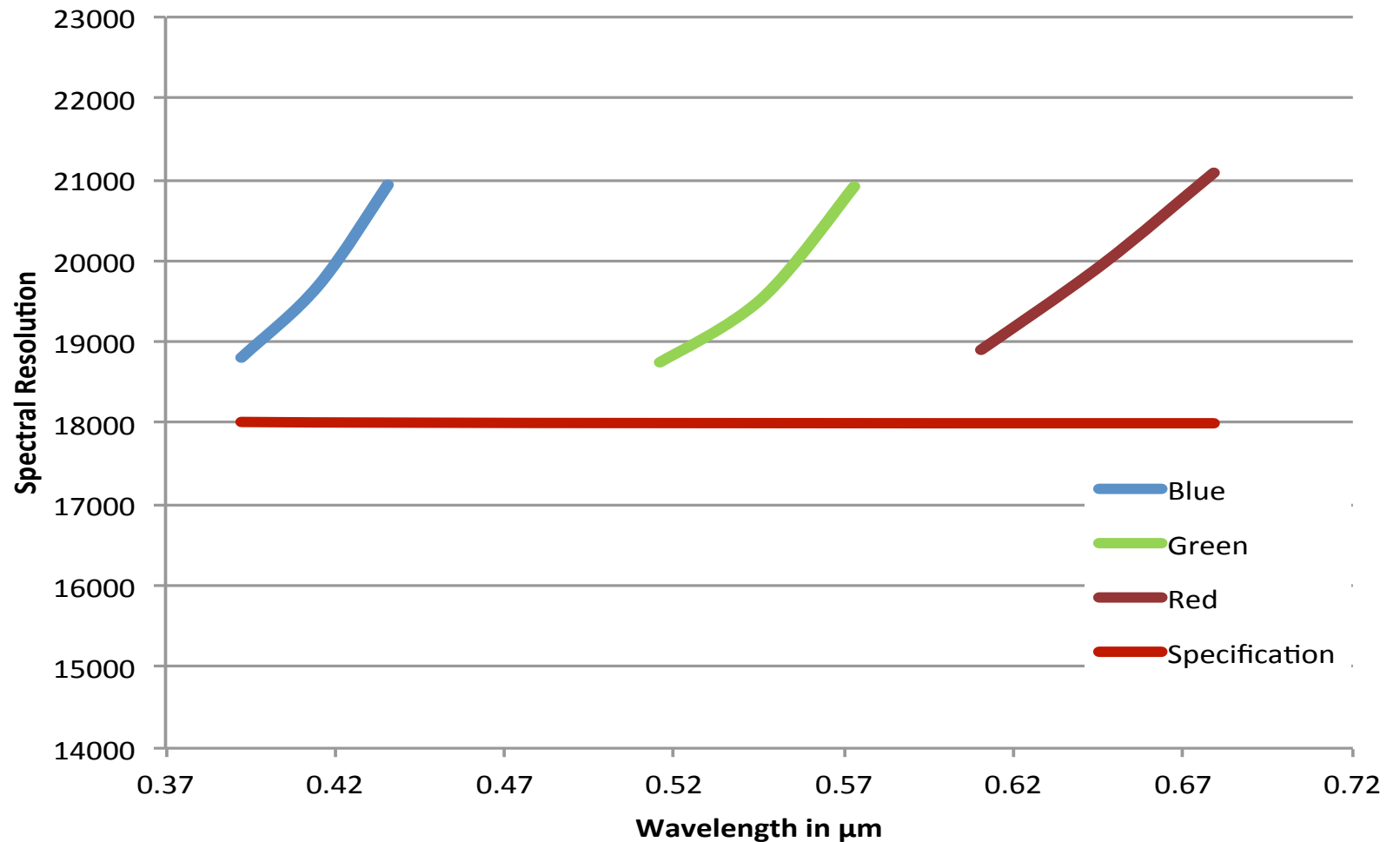


# High Resolution Spectrograph (HRS)

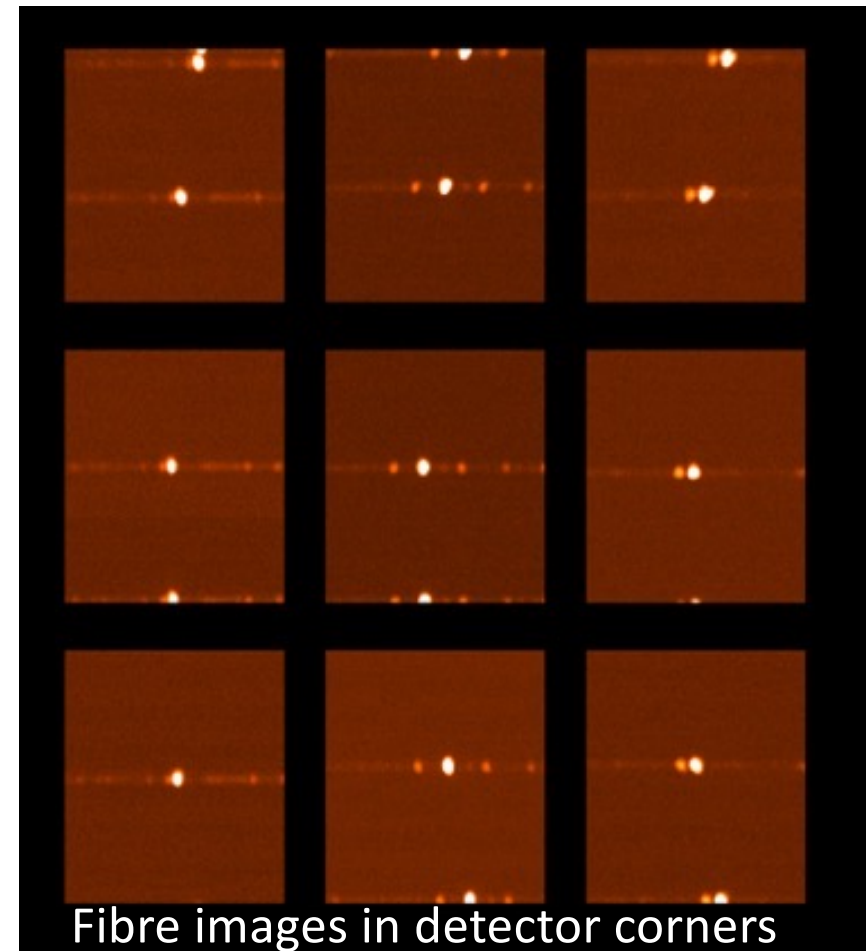
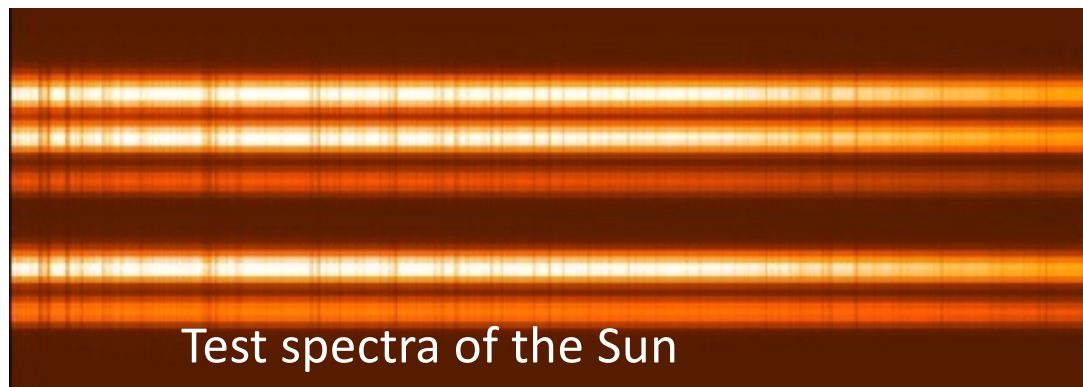
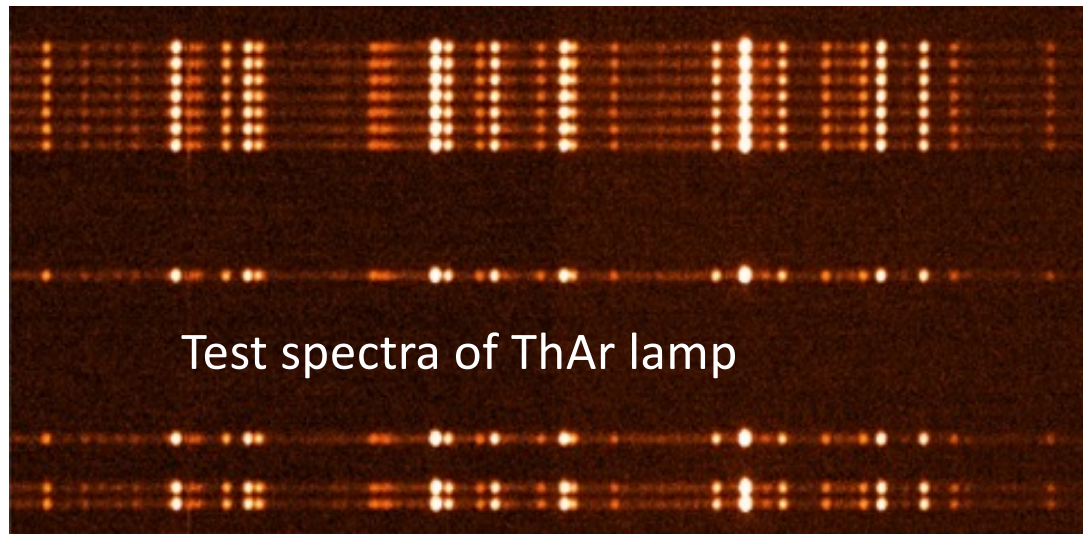


3 arms  
spectrograph  
3 CCDs 6k x 6k  
250 mm beam size  
812 science fibres  
1 spectrograph

Design and build at  
ZAH/LSW in  
Heidelberg

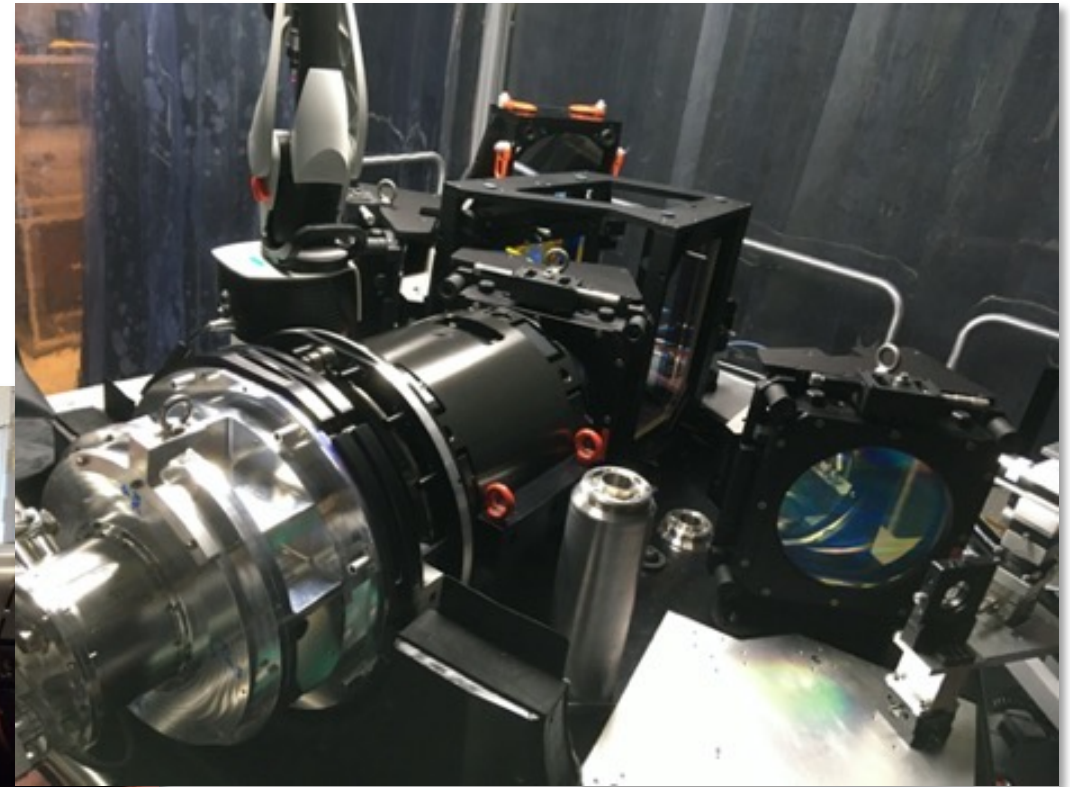
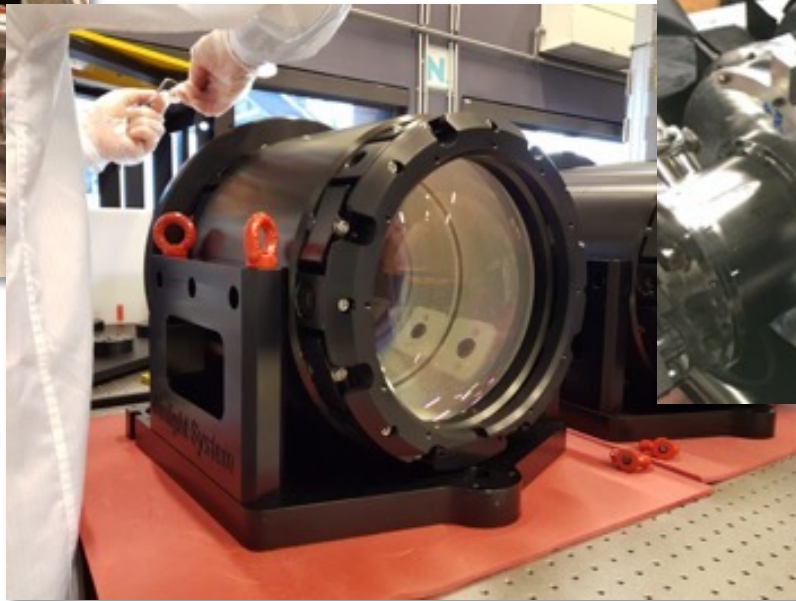
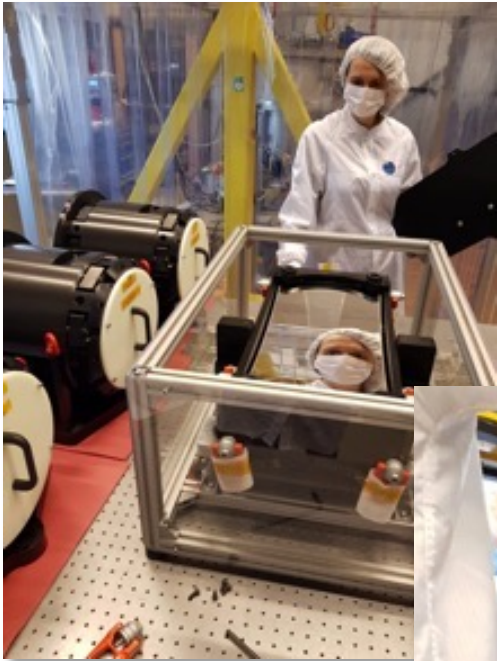


# High Resolution Spectrograph (HRS)



# Low Resolution Spectrograph (LRS)

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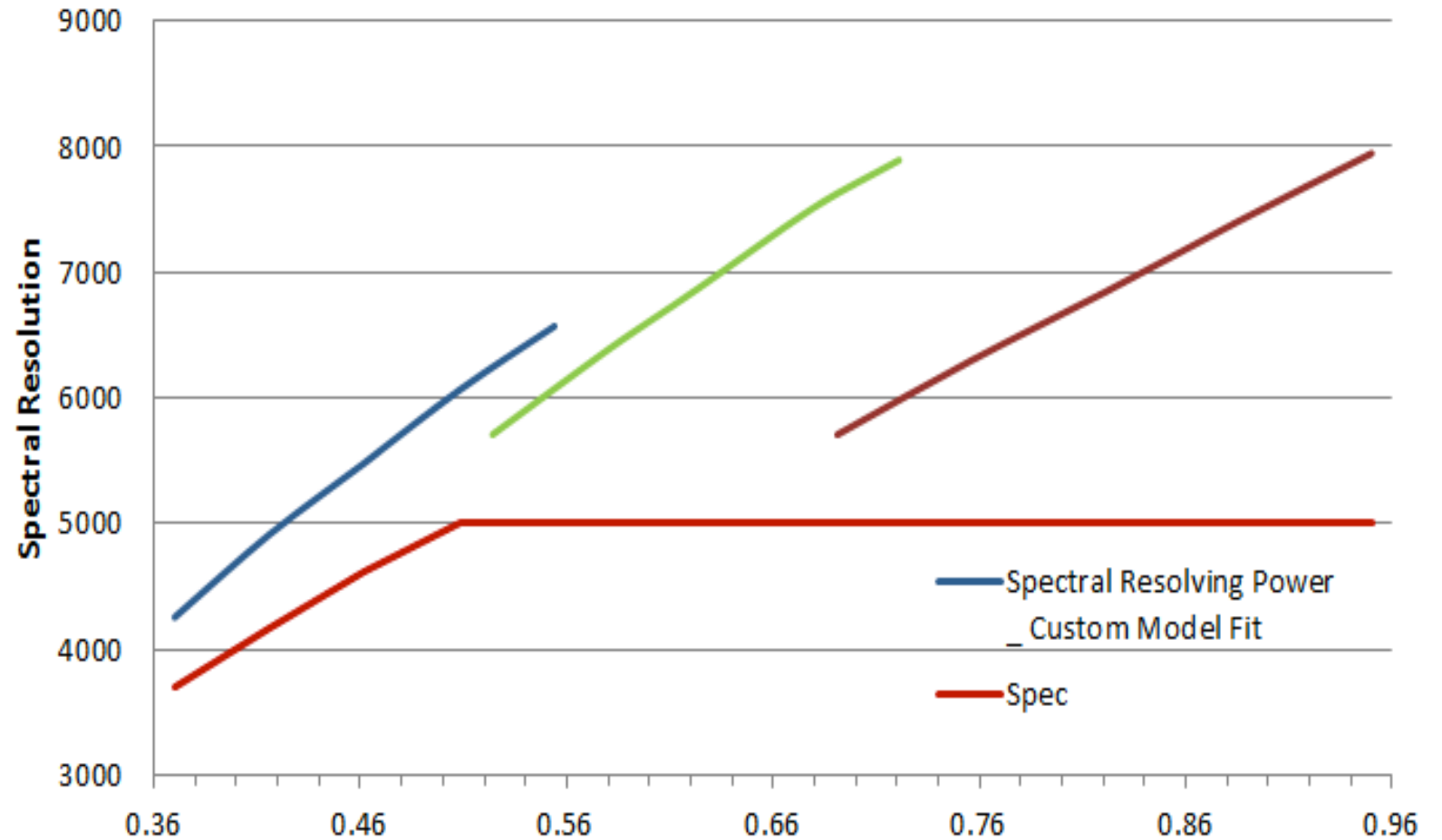




# Low Resolution Spectrograph (LRS)



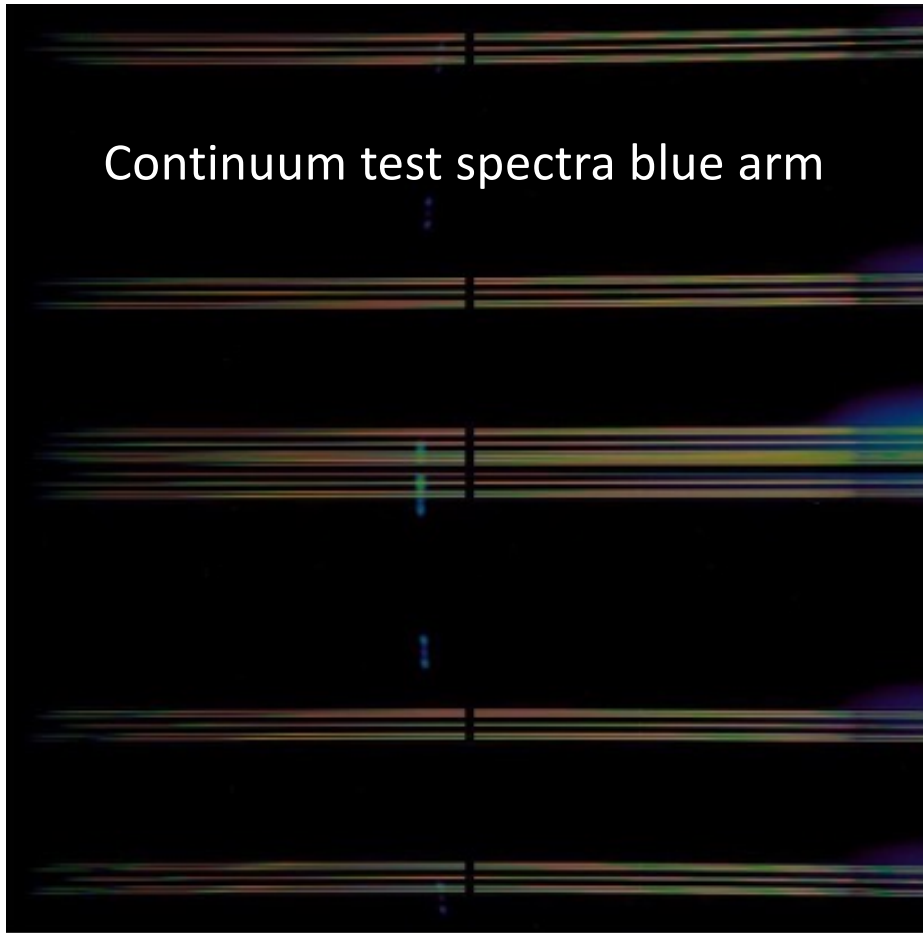
3 arms  
spectrograph  
3 CCDs 6k x 6k  
200 mm beam size  
812 science fibres  
per spectrograph  
2 mirrored  
spectrographs  
Thermally stabilized  
Design and build at  
CRAL in Lyon



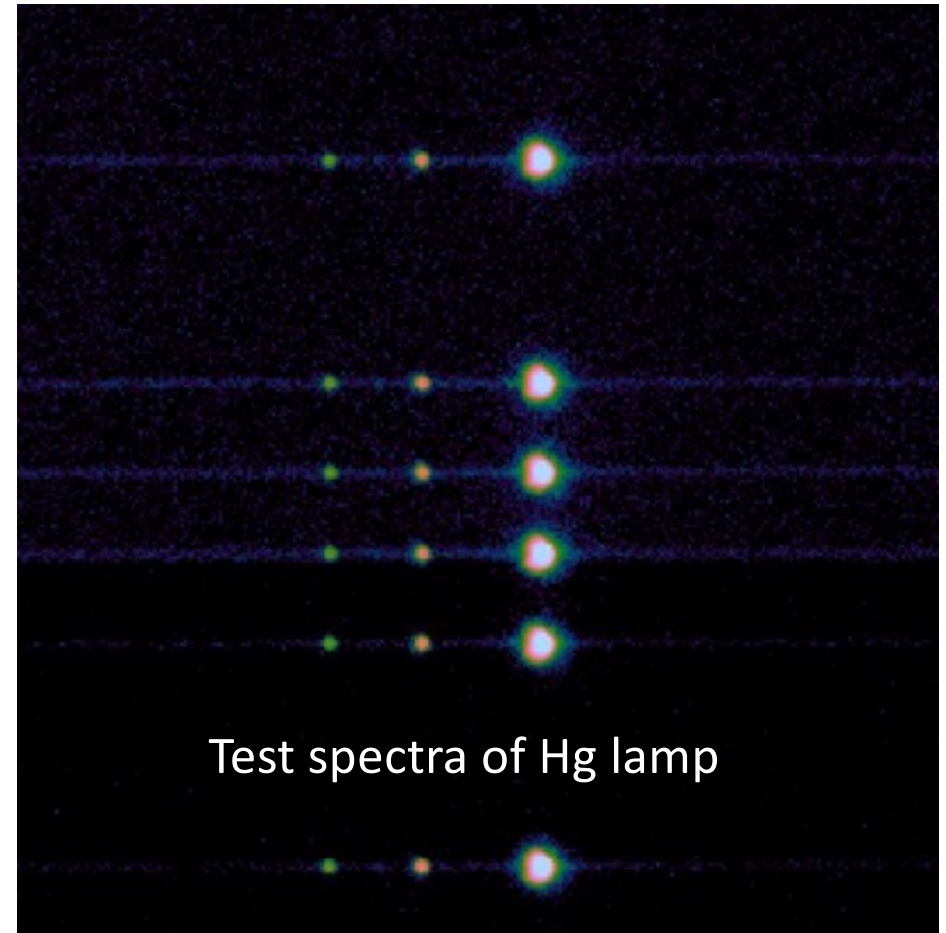
# Low Resolution Spectrograph (LRS)

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Continuum test spectra blue arm



Test spectra of Hg lamp

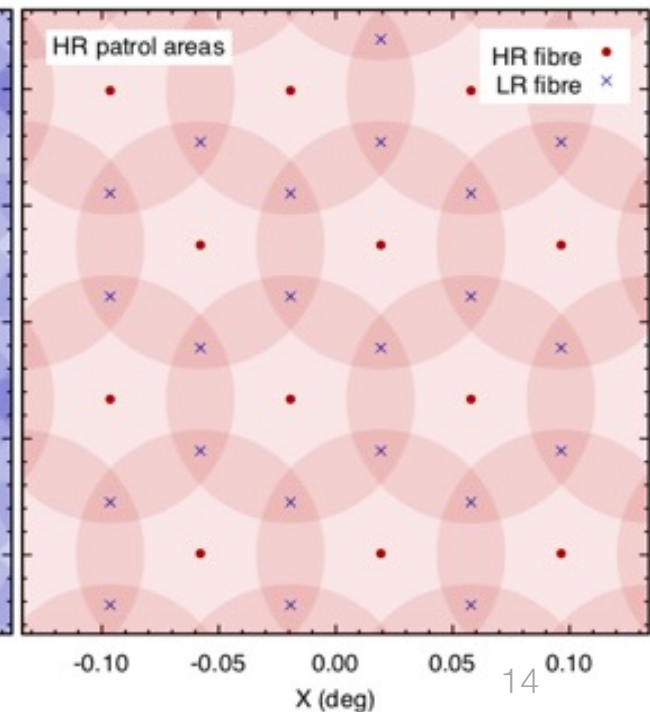
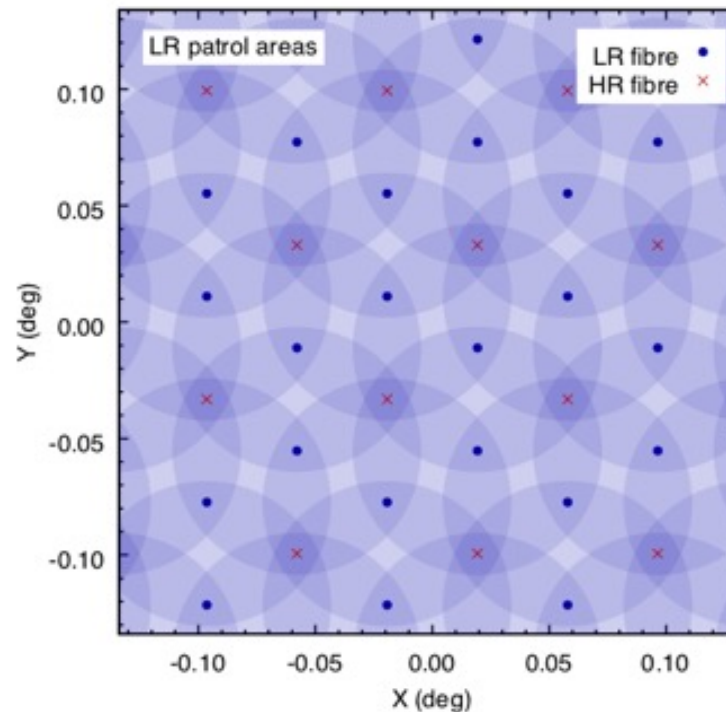
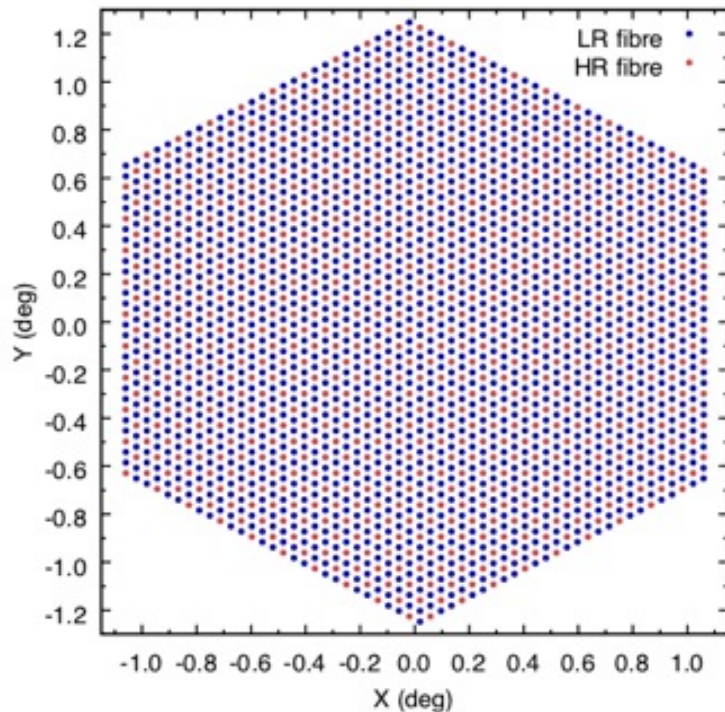
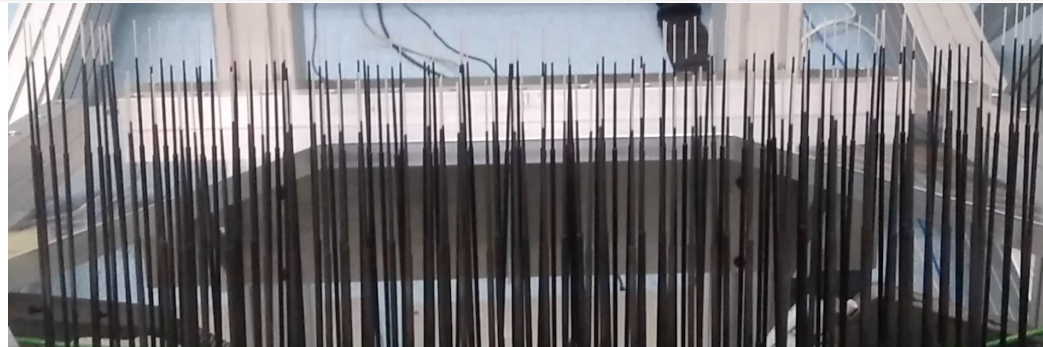


# AESOP Fibre Positioner and Metrology System



2436 Fiber Probes

- patrol diameter  
= 3.23 arcmin
- RMS accuracy <math>< 10\mu\text{m} / 0.17''</math>
- minimum separation  $\sim 20''$
- reconfiguration time <math>< 2\text{ min}</math>

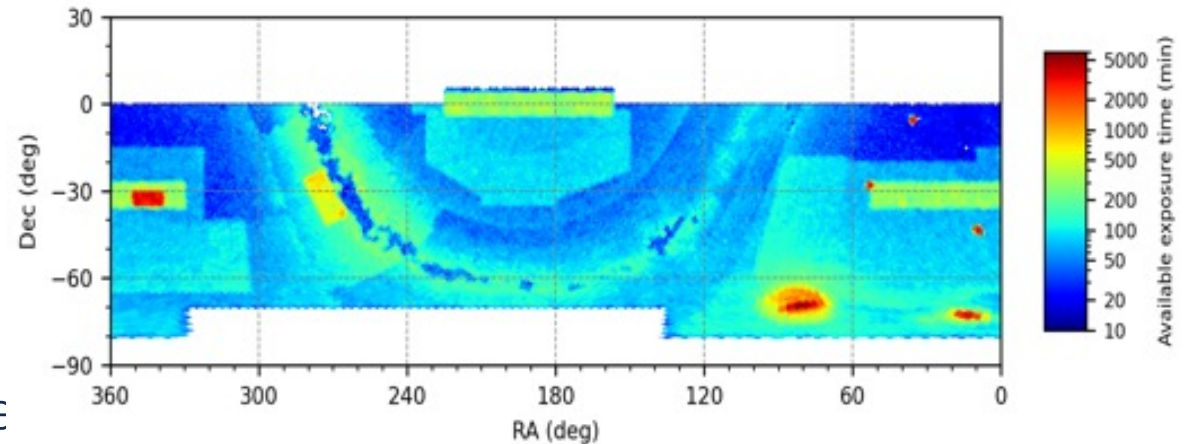




# 4MOST Operations



- Unique operations model for MOS instruments suitable *for most* science cases
- Surveys will be defined by *Consortium* and *Community*
- All Surveys will run *in parallel*
  - Surveys share fibres per exposure for increased efficiency
- *Key Surveys* set observing strategy
  - Millions of targets all sky
  - Fill all fibres
- *Add-on Surveys* for smaller surveys
  - Small fraction fibres all sky or
  - dedicated areas
  - $10^3$  to  $10^6$  targets
- Several passes of sky with exposure
- Wedding-cake distribution for total time 1h to 10h



# Policies: data and publication rights



- Data releases:
  - L0 (raw data): public immediately
  - L1 (extracted, calibrated spectra): yearly
  - L2 (derived products): negotiated by each Survey with ESO
- All data of all Surveys are accessible to all Science Team members
- Publications are limited to approved Projects. Consortium Projects may be rejected if overlapping with Community Survey Core Science and vice versa (SDSS-like policy). Also PhD Projects extra protected

# 4MOST who is who?

And 350+ others!



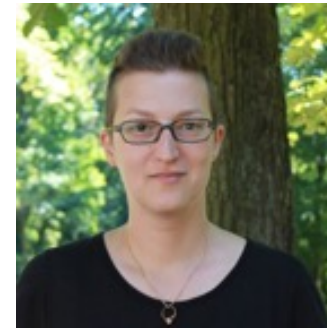
Roelof de Jong  
Principal Investigator



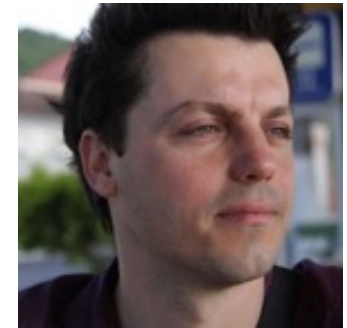
Joar Brynnel  
Project Manager



Jakob Walcher  
Operations Manager



Genoveva Micheva  
Instrument Scientist



Vincenzo Mainieri  
Project Scientist ESO



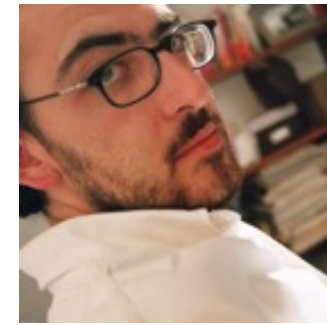
Sofia Feltzing  
Project Scientist  
Galactic



Richard McMahon  
Project Scientist  
extragalactic



Joe Liske  
Chair Science  
Coordination Board



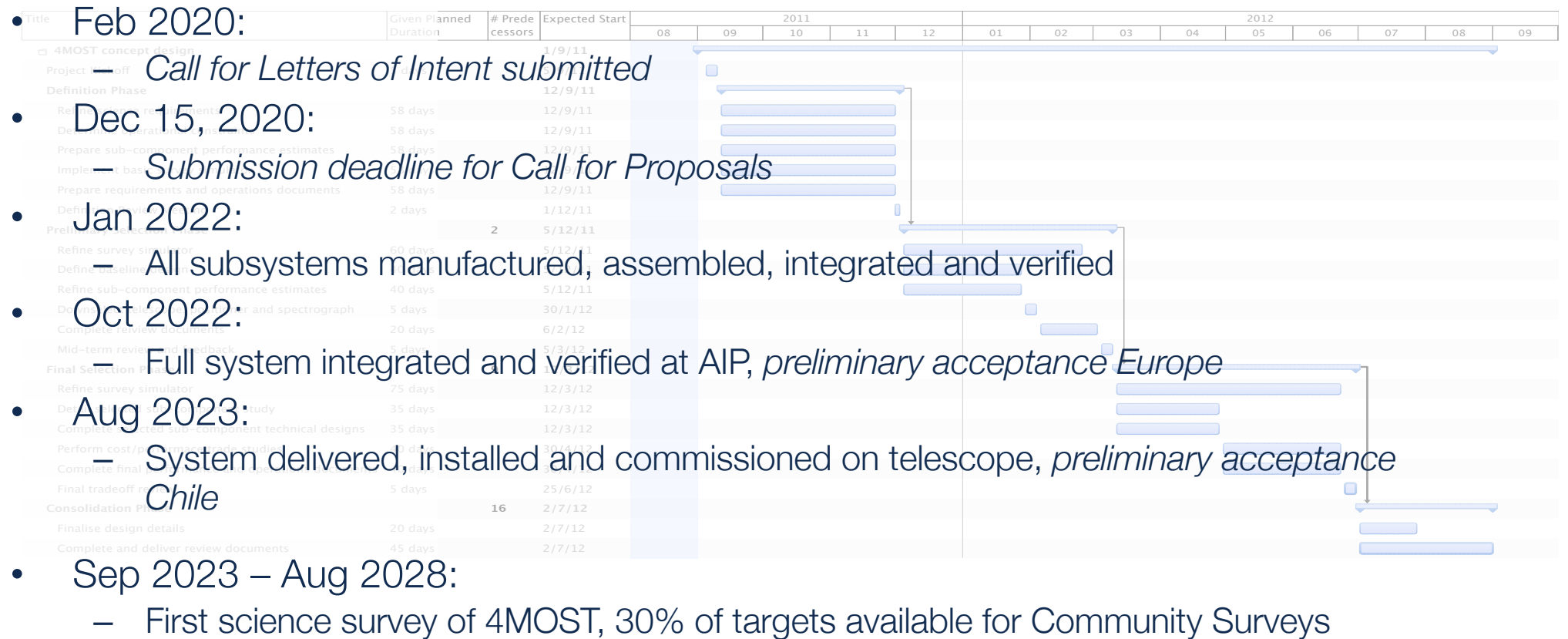
Andrea Merloni  
Lead Operations  
System



Nic Walton  
Lead Data  
Management



# Schedule and Milestones





# Wide-field, high-multiplex optical spectroscopic survey facility for ESO



[www.4MOST.eu](http://www.4MOST.eu)

- Status:
  - FDR completed, in construction phase, *operations start Q3-2023* ( $\geq 2 \times 5$  year)
  - Call for Proposals 15 Dec 2020!
- Science:
  - *Cosmology, galaxy evolution, high-energy and Galactic science*
  - Complement large area space missions: Gaia, eROSITA, Euclid, PLATO
  - Complement ground-based surveys: VISTA, VST, DES, LSST, SKA, etc.
- Survey facility:
  - Instrument, science operations, data products, science
  - Run all-sky 5 year *public* surveys in parallel with yearly data releases
  - Key surveys organized by consortium, add-on surveys from community through ESO
- Instrument specifications:
  - High multiplex: 1600 fibres @  $R \sim 6000$  + 800 fibres @  $R \sim 20,000$  in parallel
  - Wavelength: LR: 370-950 nm HR: 392-437 & 515-572 & 605-675 nm
  - Large field-of-view on VISTA, 4m-class telescope:  $\varnothing = 2.6^\circ$