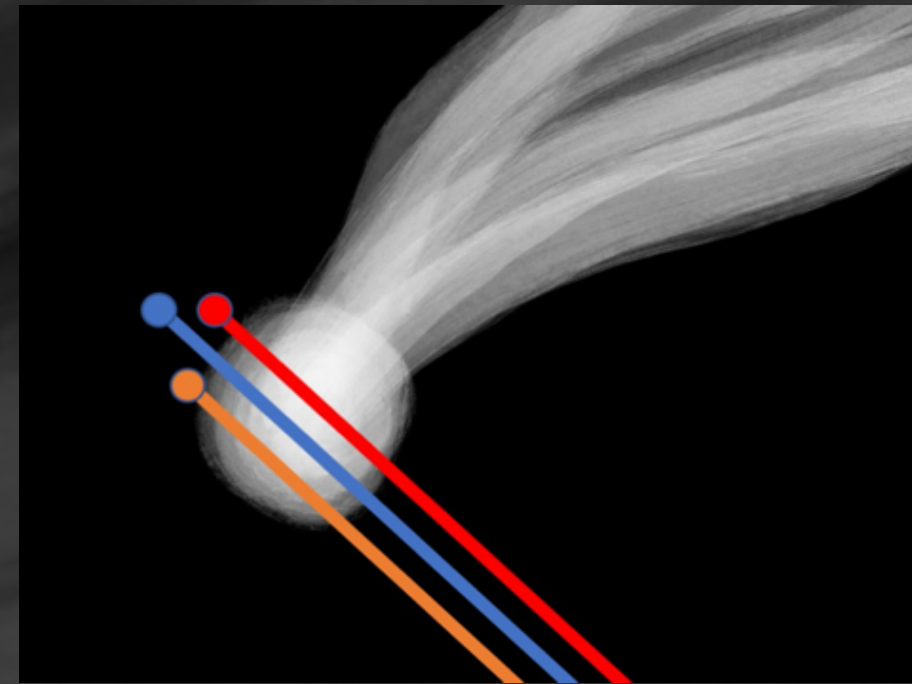
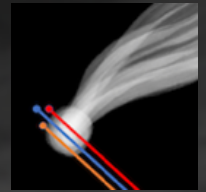


# The importance of LSST to ESA's Comet Interceptor mission



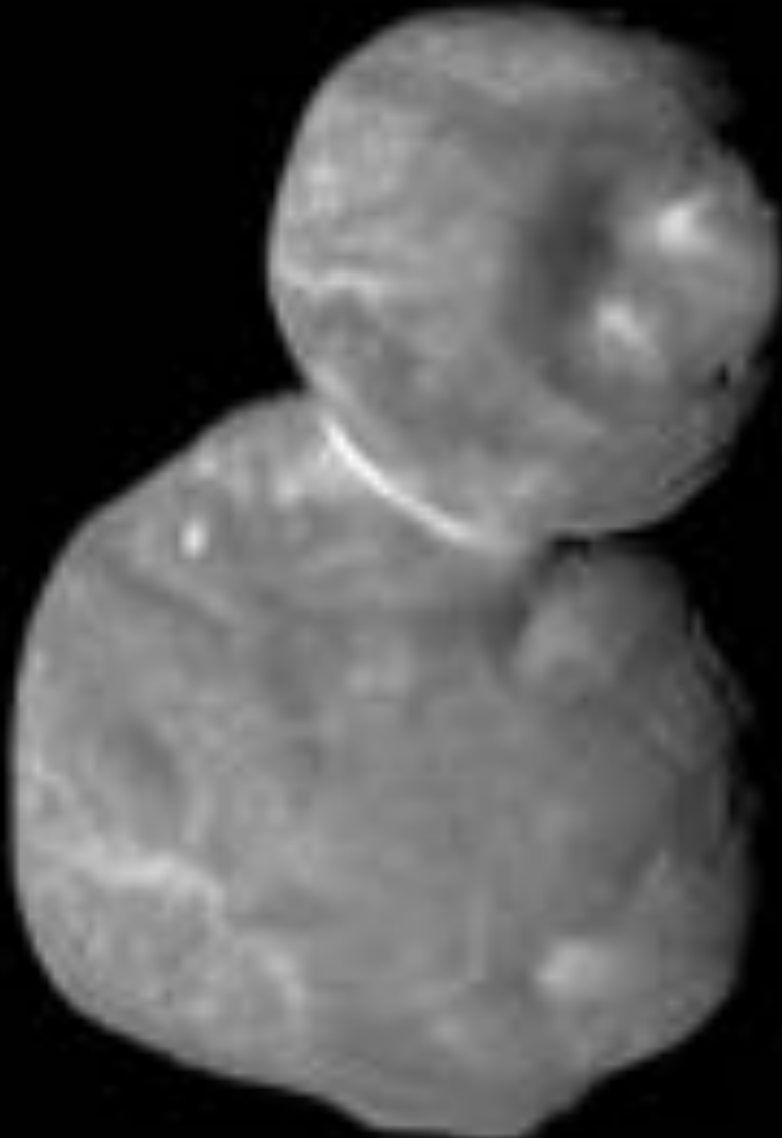
Comet Interceptor is a mission targeting a long-period comet, preferably dynamically-new, or an interstellar object.



## Why?

- All previous comet missions have been to objects that have passed the Sun many times
- Those comets have changed over time, and are covered in a thick layer of dust
- A dynamically-new comet is one that is probably nearing the Sun for the first time
- These are **pristine**





1P/Halley  
16 × 8 × 8 km  
Vega 2, 1986



81P/Wild 2  
5.5 × 4.0 × 3.3 km  
Stardust, 2004



67P/Churyumov-Gerasimenko  
4 × 3 km  
Rosetta, 2014



103P/Hartley 2  
2.2 × 0.5 km  
Deep Impact/EPOXI, 2010



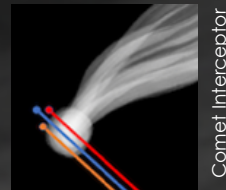
19P/Borrelly  
8 × 4 km  
Deep Space 1, 2001



9P/Tempel 1  
7.6 × 4.9 km  
Deep Impact, 2005



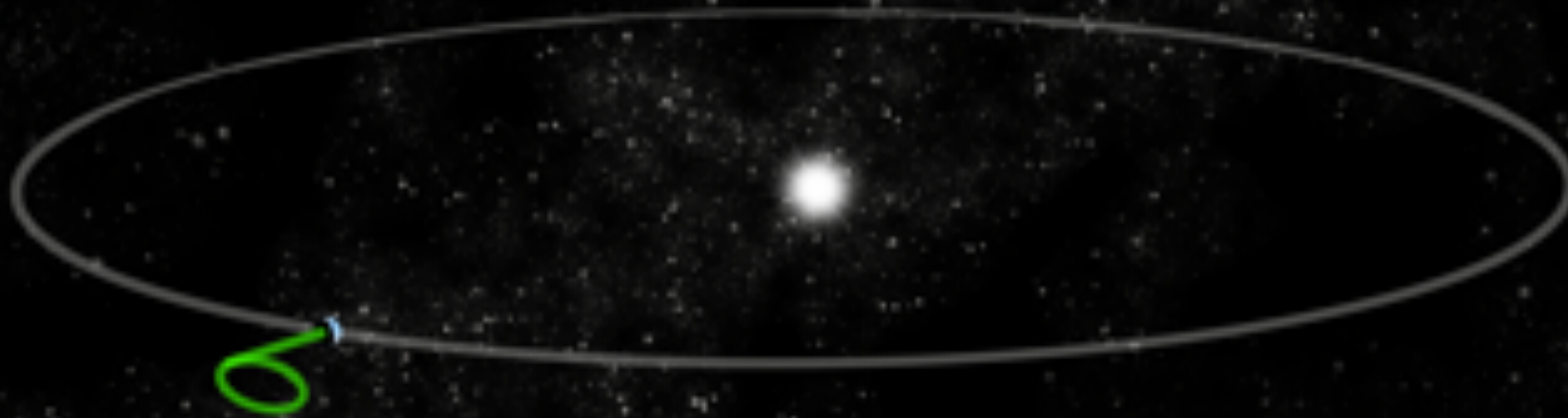
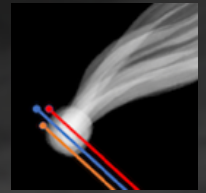
ESA / Rosetta / MPS for OSIRIS Team;  
MPS/UPD/LAM/IAA/SSO/INTA/UPM/DASP/IDA



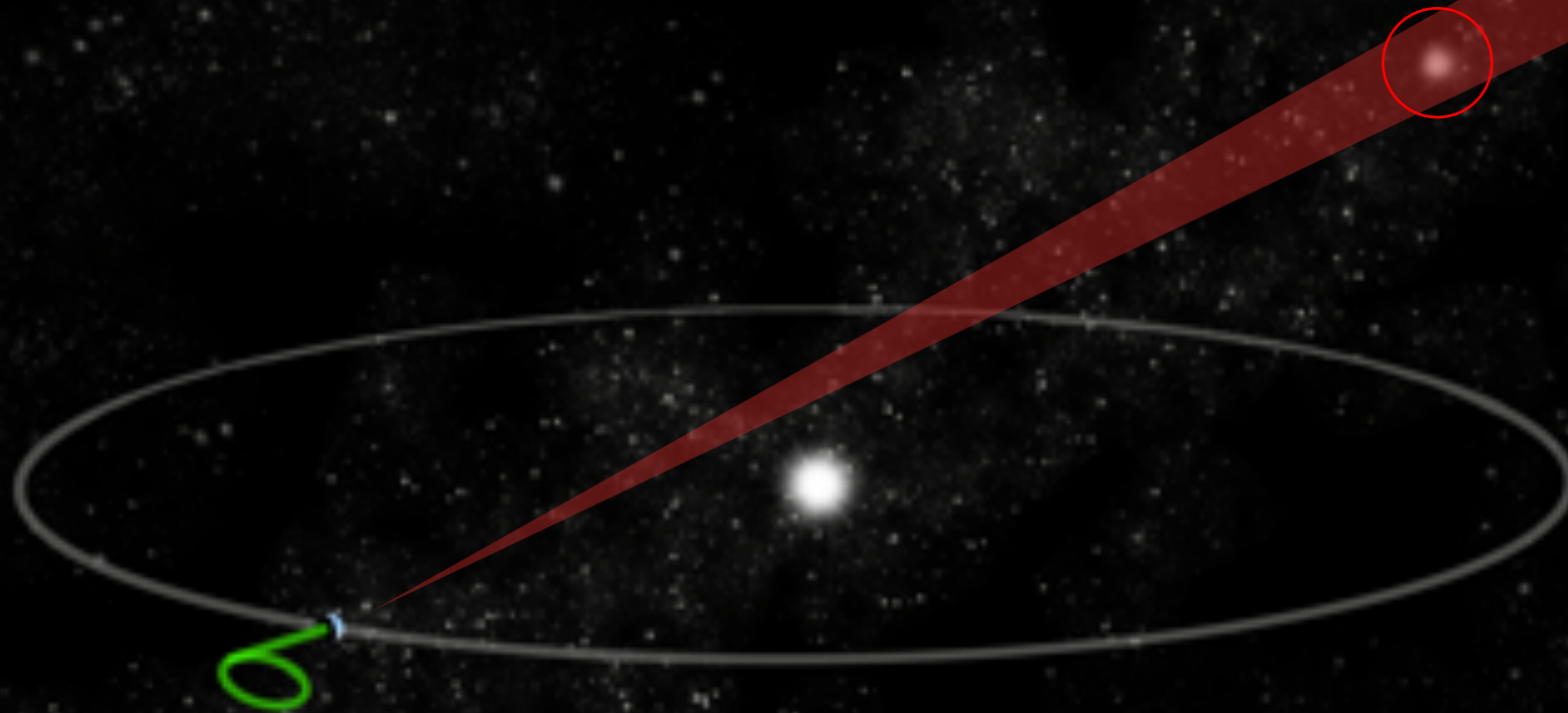
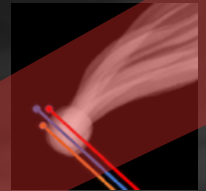
# Comet Interceptor

- New comets are unknown targets, can't plan and build mission years in advance
- Need to build a flexible mission and pick the target later
- ~few years warning, best way is to launch and wait in space
- ESA's first F-class mission gives an opportunity to do this:
  - Shared launch with Ariel to Sun-Earth L2 point
  - Wait
  - Set off on intercept course when comet discovered
  - Fast fly-by reconnaissance – first view of a new comet

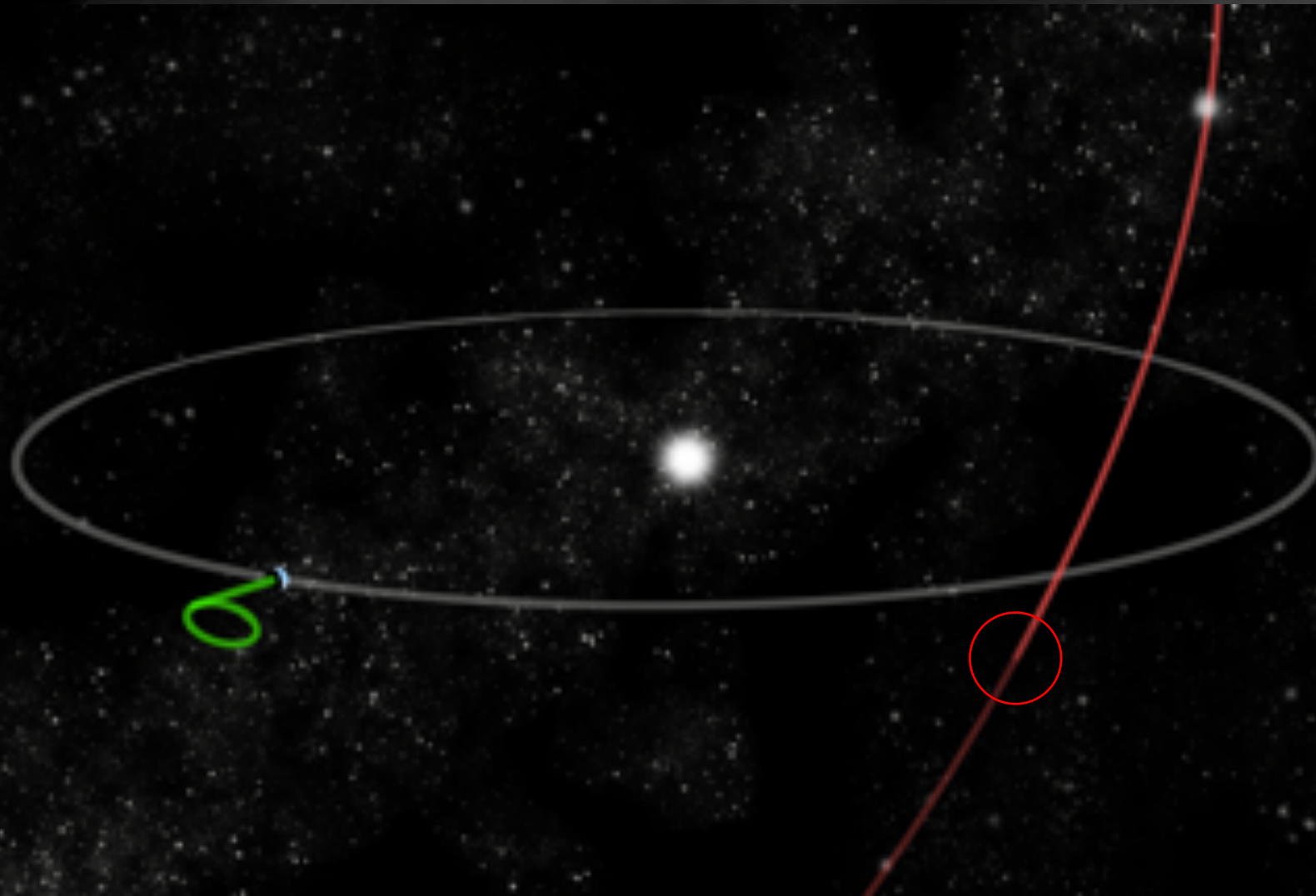
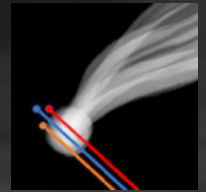
- Mission 'parked' at stable Lagrange point L2 after launch with Ariel
- Waits for up to 2-3 years for new target discovery



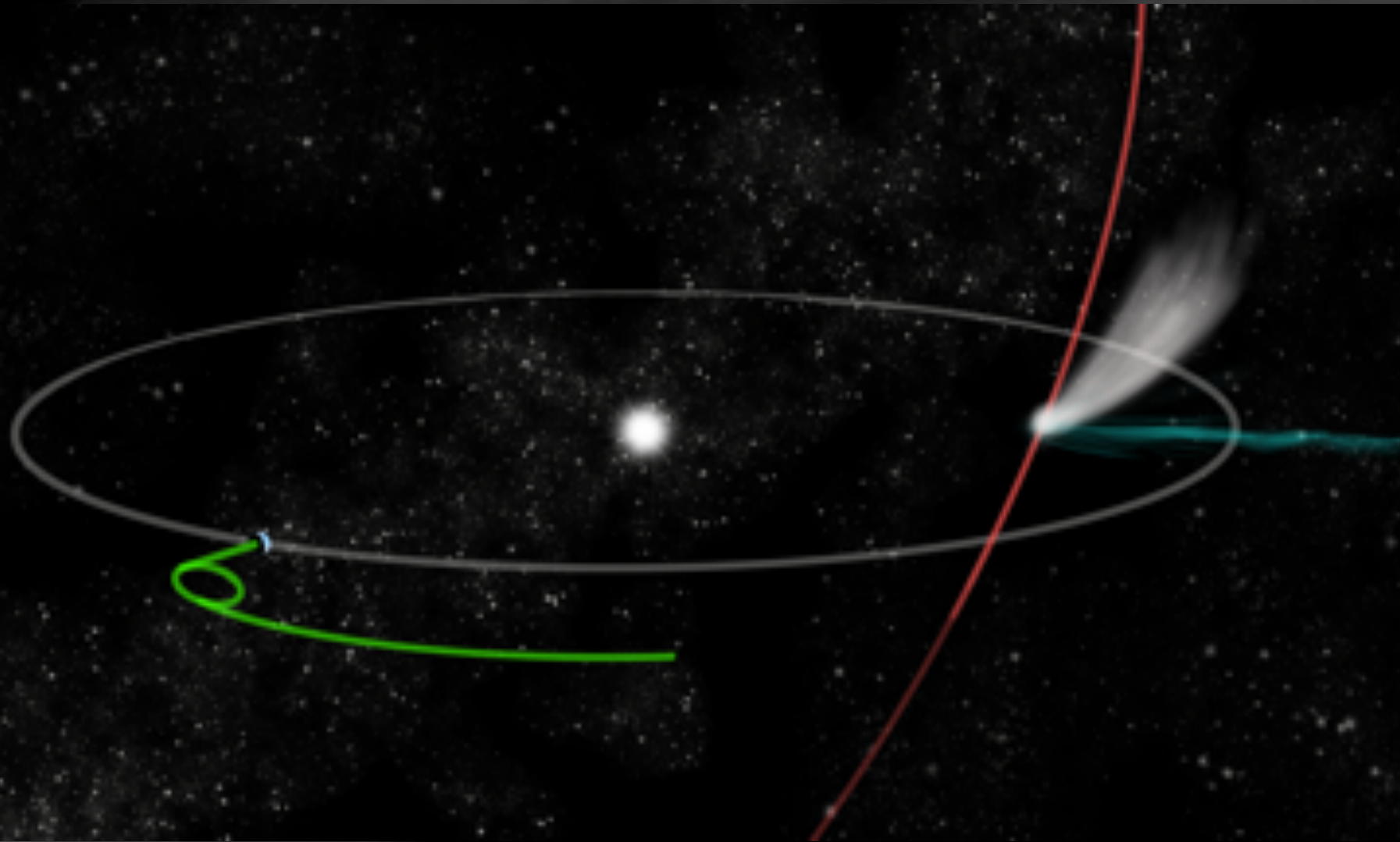
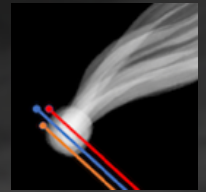
- Target discovered by a ground-based observatory (probably Rubin)



- Orbit computed and ecliptic crossing point predicted

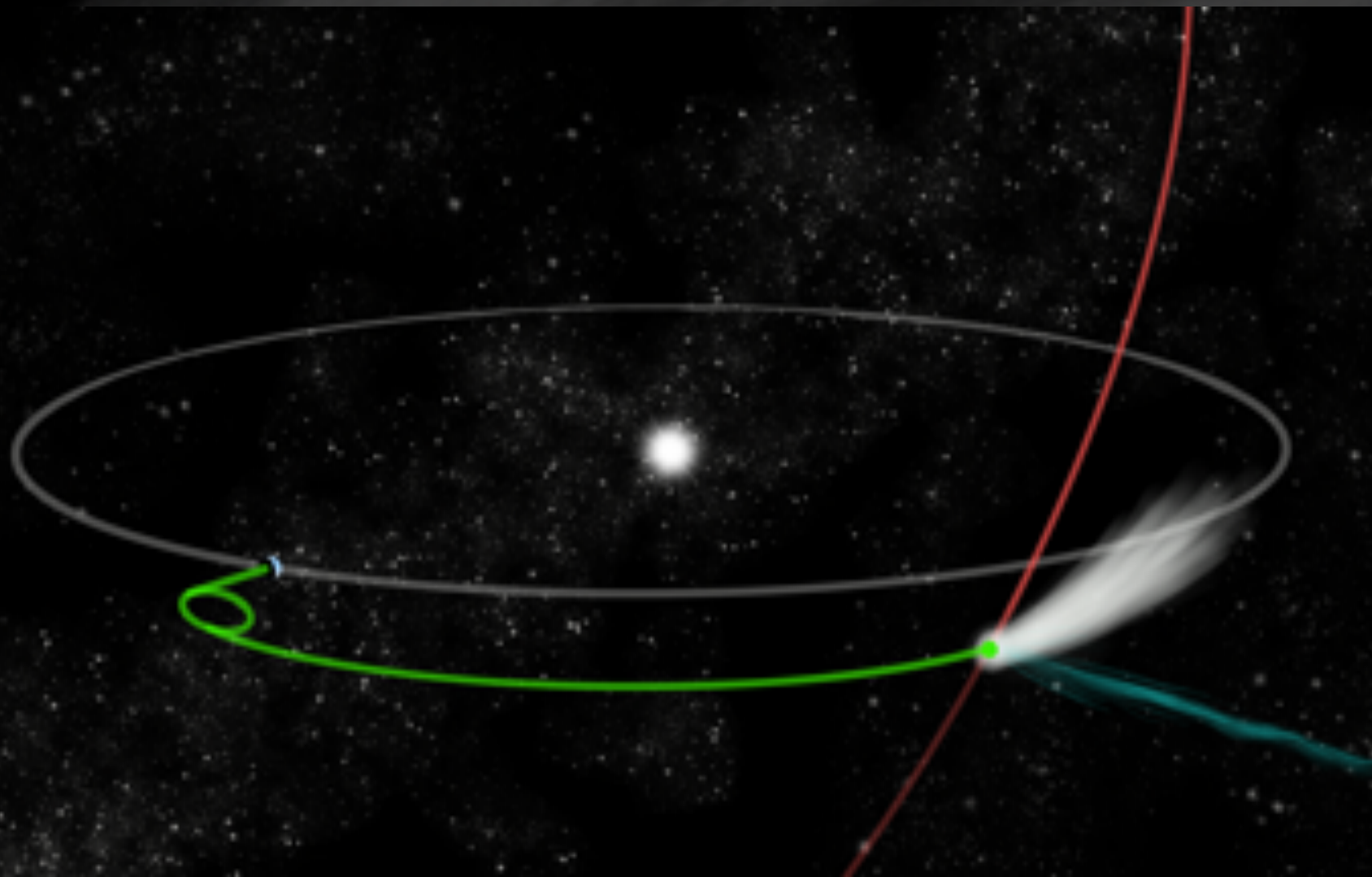
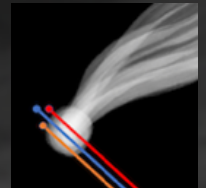


- Comet Interceptor leaves L2 to intercept comet's path

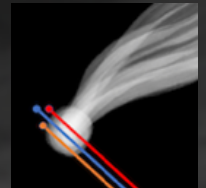




- Encounter with comet close to the ecliptic plane



# A Multi-Spacecraft Mission



- A: main spacecraft

safe / distant measurements



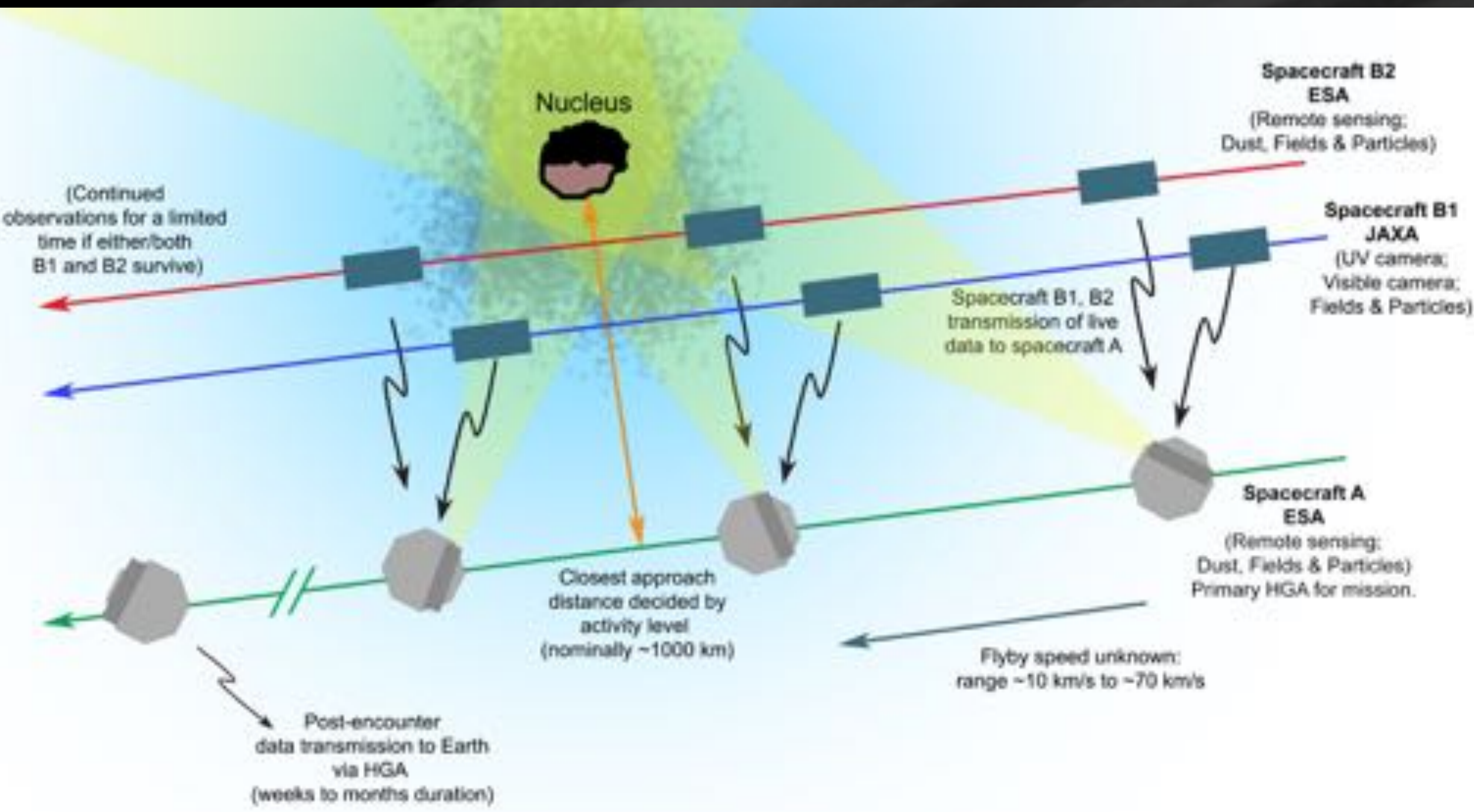
- B1: inner coma



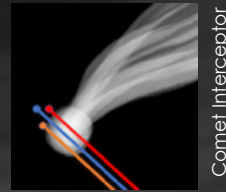
higher risk / high gain  
closer approaches  
to nucleus



- B2: nucleus + coma



# Example – C/2001 Q4 (NEAT)



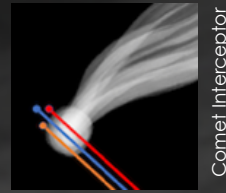
- Based on previous bright dynamically-new comet
- Real comet found ~3 years out
- VRO-LSST would have found it ~8 years out
- Target would have been known before launch
- ~1.5 year wait
- ~3 year cruise

*Table 7.2.1. Dates of key events in example mission to C/2001 Q4 (NEAT).*

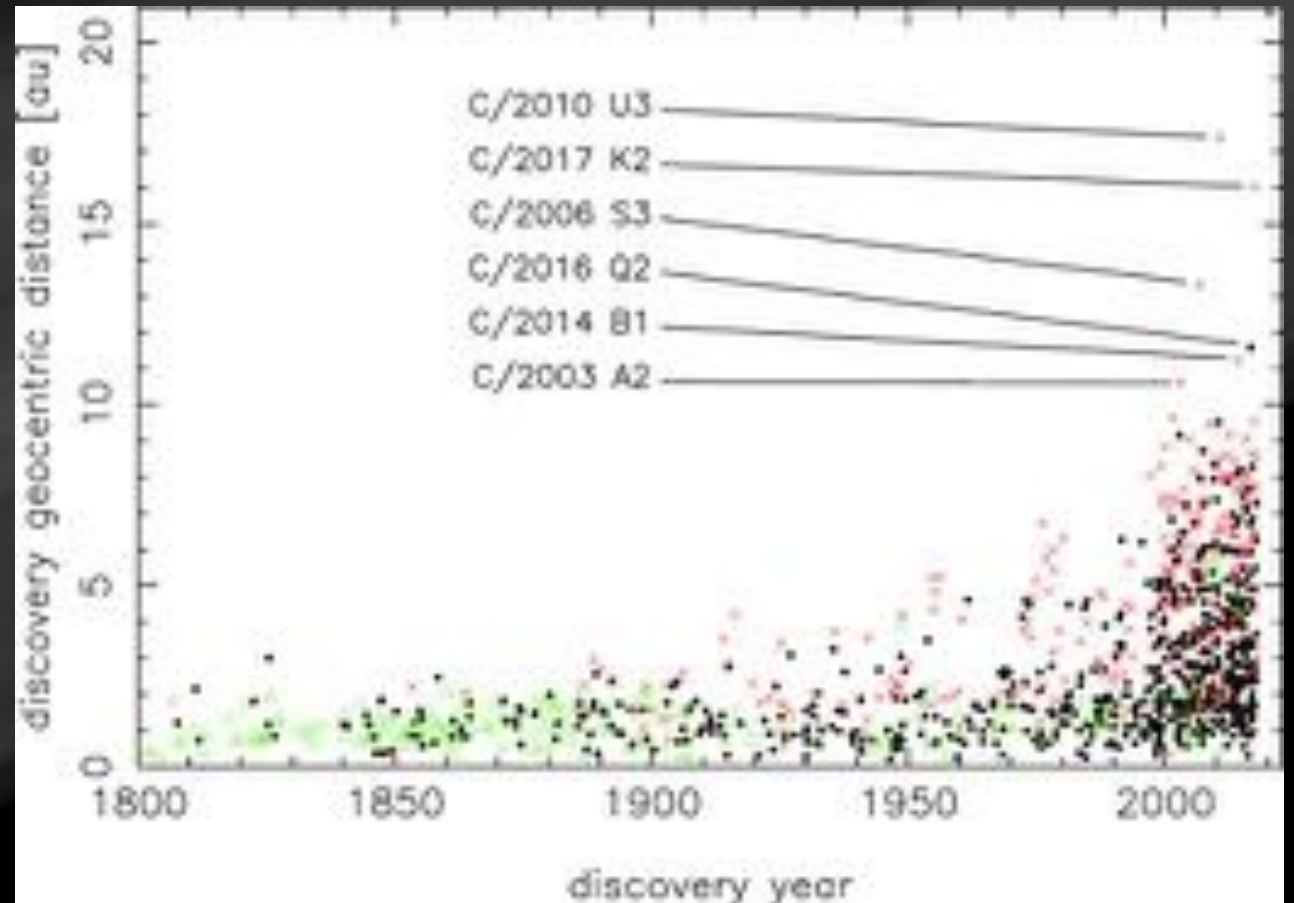
Event	Date
LSST discovery	~July 1996
Launch	10-Dec-1999
Departure from L2	29-Jul-2001
<i>Real discovery</i>	<i>24-Aug-2001</i>
OP Nav images begin	Jan 2004
Flyby	14-May-2004
End of mission	Nov 2004



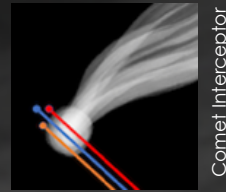
# Discovery of distant comets



- Distance at which comets are discovered is increasing
- Comets have shown activity at 20-40 au in deep images.
- With LSST, expectation is that discovery at large distances will be routine
- Gives warning times of years
- Enables trajectory optimization to give mission with OK wait times and fuel mass



# Comet Interceptor trajectories



- Longer warning time means a greater range of comets can be reached
- Increases the chances of a successful mission to a new comet

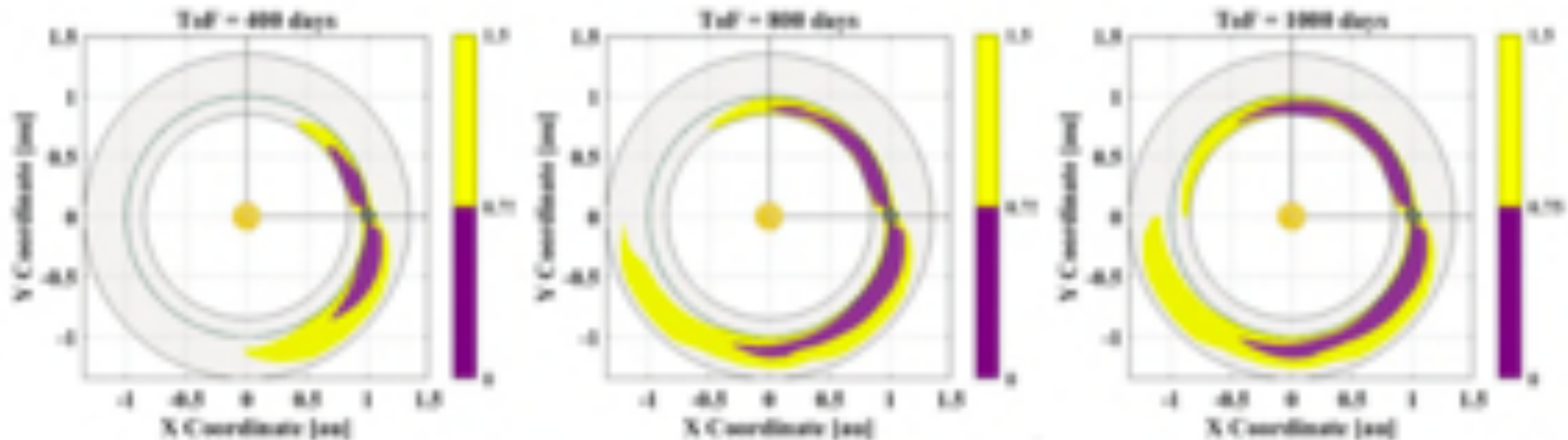
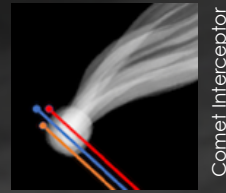


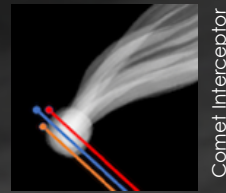
Figure 5c: Accessible regions for chemical propulsion. The color map represent the  $\Delta v$  budget in km/s

# LSST simulations



- Critical for detailed assessment of CI success is the question of how far from the Sun LSST will find comets, and at what rate
- New survey simulator tool being developed (see flash talk by Grigori Fedorets) that will be used to predict comet discovery rate
- Inputs:
  - Distribution of expected comet orbits
  - Comet brightness model
  - Input distribution of nucleus sizes, peak activity/brightness and slope parameter

# LSST simulations



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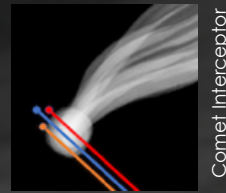
## - Inputs:

Set of 1700 synthetic comets from Boe et al 2019,  
Based on model by Wiegert & Tremaine 1999

- Distribution of expected comet orbits
- Comet brightness model
- Input distribution of nucleus sizes, peak activity/brightness and slope parameter



# LSST simulations



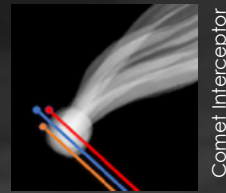
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Simple brightness model:

$$A_f p = A_f p_1 \times r^k$$

$$k = -4 \text{ (SPC)}, -2 \text{ (LPC)} \text{ ?}$$

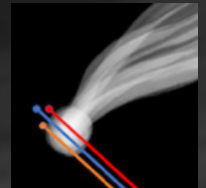
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  - Comet brightness model
  - Input distribution of nucleus sizes, peak activity/brightness and slope parameter

Based on historic comet observations

# LSST simulations



- Critical for detailed assessment of CI success in far from the Sun LSST will find comets, and
- New survey simulator tool being (developed by Fedorets) that will be used to simulate LSST observations

## Inputs:

- Distribution

- 

... sizes, peak activity/brightness and slope

**COMING SOON! Simulator nearly complete, aim to have simulations done this summer to feed into CI system SRR in October**

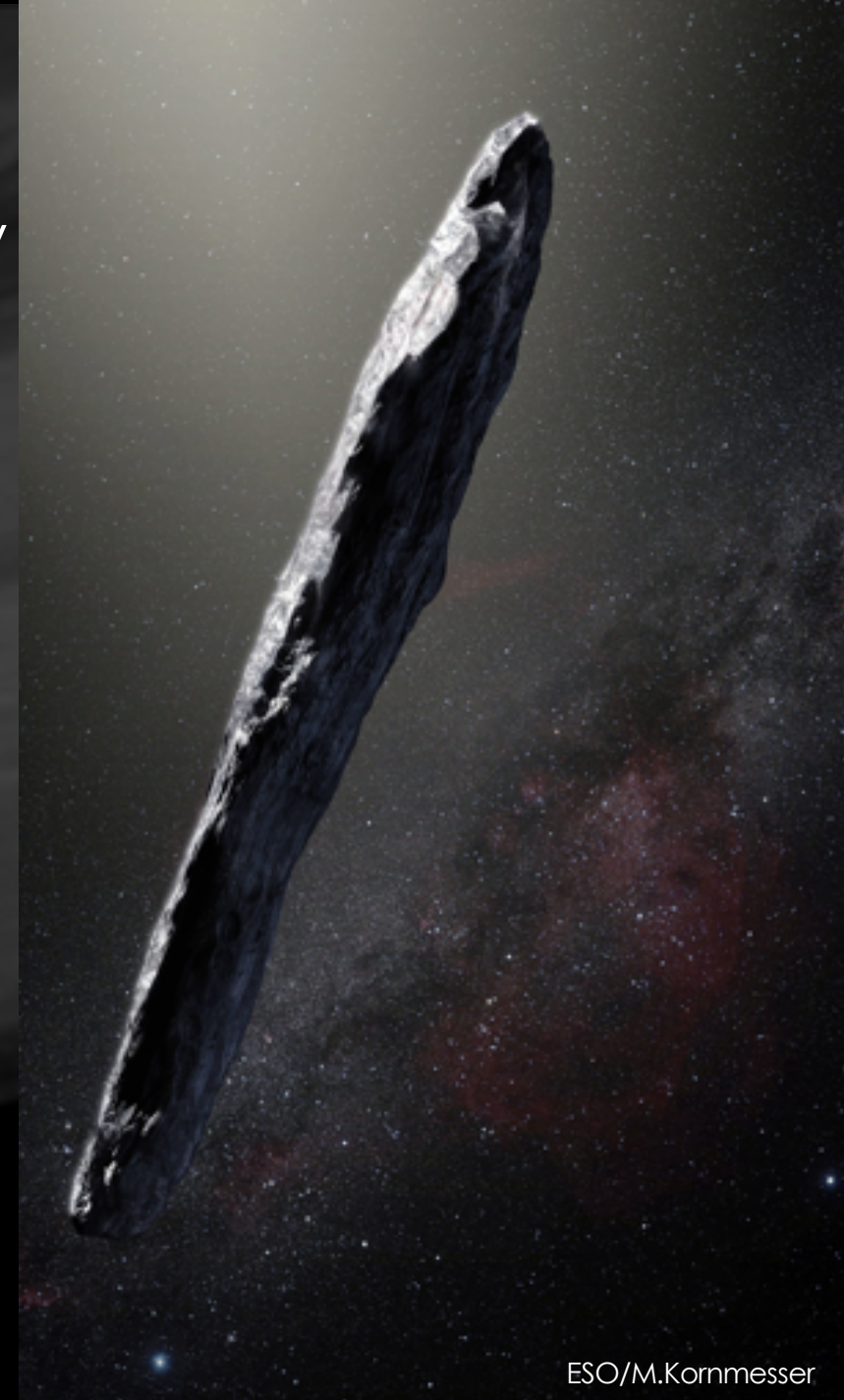
**Critical to demonstrate that LSST will do what we need it to for CI mission adoption in 2022**

# Interstellar targets?

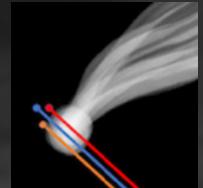
- 'Oumuamua study\* showed that Rubin Observatory could find one accessible target in ~10 years.
- A small but non-zero chance of a suitable target within 2-3 years
- Comet 2I/Borisov – a sign of promising discovery statistics?



NASA



ESO/M.Kornmesser



Comet Interceptor



<http://www.cometinterceptor.space/>



@cometintercept